



## AIRS Forward Model Validation and Status

AIRS Science Team Meeting: Nov/Dec 2004



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Estimated *AIRS* RTA accuracy via ARM-TWP and ECMWF validation studies.

RTA accuracy now on order of instrument accuracy (except for high-altitude water and Non-LTE). Maybe another factor of 2-3x improvements to reach instrument relative accuracy.

RTA accuracy in upper troposphere, stratosphere hard to validate.

*AIRS* sees variability in  $\text{CO}_2$ ,  $\text{CO}$ ,  $\text{SO}_2$ ,  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ . No  $\text{N}_2\text{O}$  or  $\text{SO}_2$  in standard RTA.

Look more carefully at *AIRS* spectral calibration for climate studies.

Do cloud-cleared data show same bias characteristics? (Wednesday)

Preliminary work with SARTA-Scattering shows reasonable abilities to simulate dust and cirrus. Particle habit, dust indices of refraction, aerosol altitude, as always, present challenges. (Thursday)

## Climate with AIRS

- Is the DAAC record for weather or climate? I assume climate.
- Climate requirements allow higher standard deviations, but lower mean errors.
- Need L1b, RTA to track instrument calibration changes and slow atmospheric changes ( $\text{CO}_2$ )
- AIRS has additional climate information:
  - IR dust forcing
  - IR cirrus (thin)
  - Minor gases ( $\text{CO}$ ,  $\text{CH}_4$ ,  $\text{CO}_2$ ,  $\text{SO}_2$ , maybe  $\text{N}_2\text{O}$ )
  - Surface emissivity
- Level 1b may be most important climate record
  - How inform users of subtle instrument changes in L1b?
    - Frequency calibration
    - Fringes

## RTA Liens (over mission)

1. (Lev 1b:) Frequency calibration (Level 1b or RTA):  $\pm 0.1\text{K}$  max
2. (Lev 1b:) Fringes (Level 1b or RTA):  $\pm 0.3\text{K}$  max
3. (Lev 1b:) Scan asymmetry:  $0.1\text{K}$  max, surface channels only
4. (Lev 2:) Cloud-cleared radiance accuracy (Wednesday)
5. Spectroscopy:  $0.2\text{K}+$ ? (upper trop/strat not validated),  $6\text{K}$  (non-LTE)
6. Parameterization accuracy: generally  $< 0.05\text{K}$
7. Regression profiles sufficiently diverse? ??
8. Variable gases:  $\text{N}_2\text{O}$ :  $0.7\text{K}$ ,  $\text{CO}_2$ :  $0.8\text{K}$ ,  $\text{CH}_4$ : ?,  $\text{SO}_2$  and  $\text{CO}$  even more
9. Use of RTA above cloud deck: ??
10. Reflected thermal for low emissivity land scenes:  $0.5\text{K}$  or more
11. Dust:  $5\text{K}+$  (makes it through cloud clearing) (Thursday)
12. Cirrus:  $\text{N/A}$  (Thursday?)
13. Emissivity variations with SST:  $0.3\text{K}$

Note: Bias stability may be  $< 0.01\text{K}$  per year!

Would like RTA stability to approach this number??

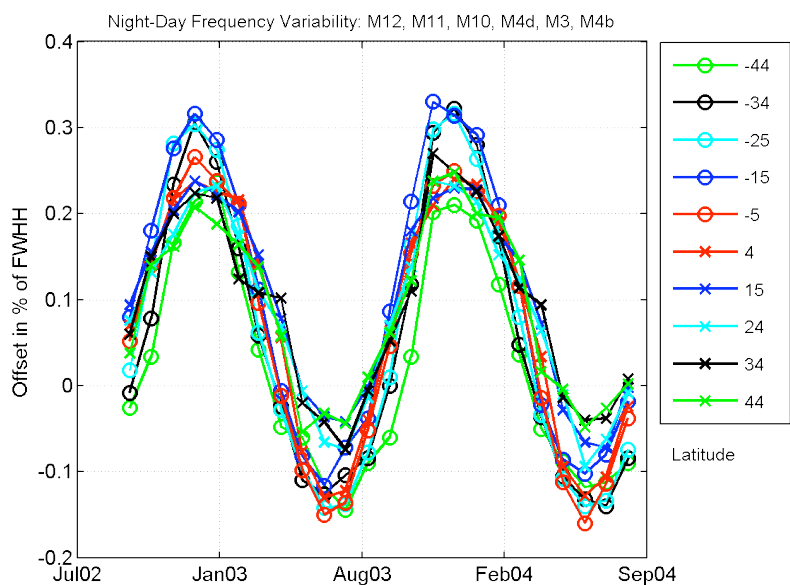
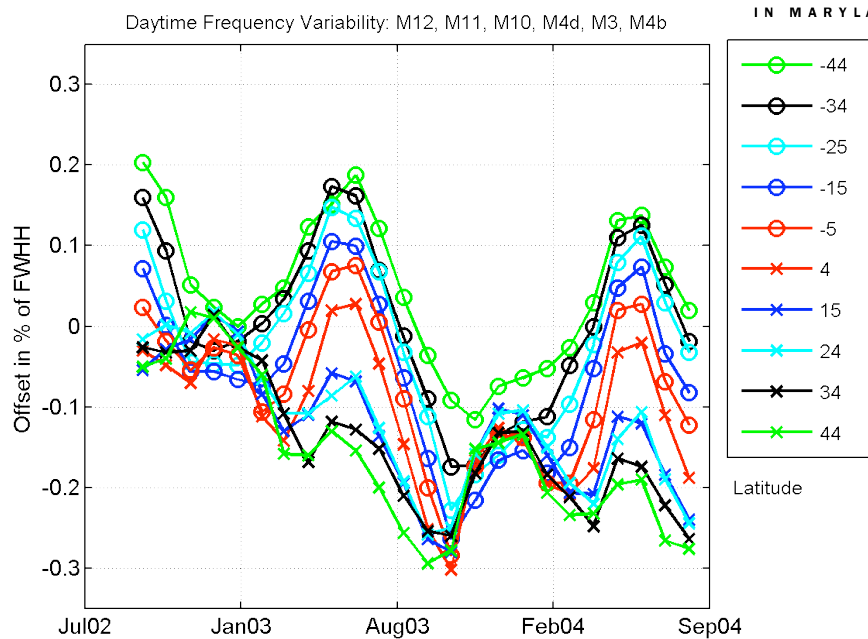
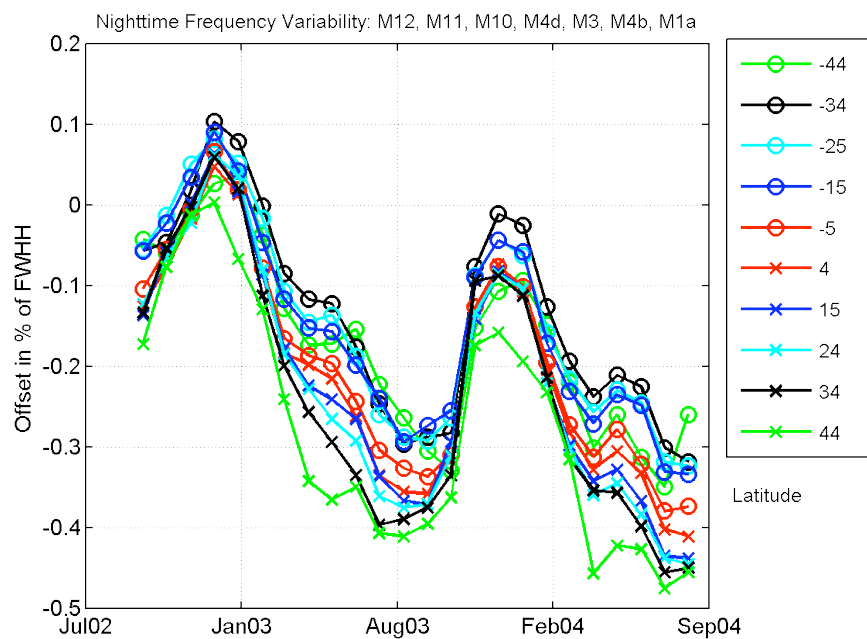
Being worked

Worked in past

Difficult problem

# Frequency Calibration

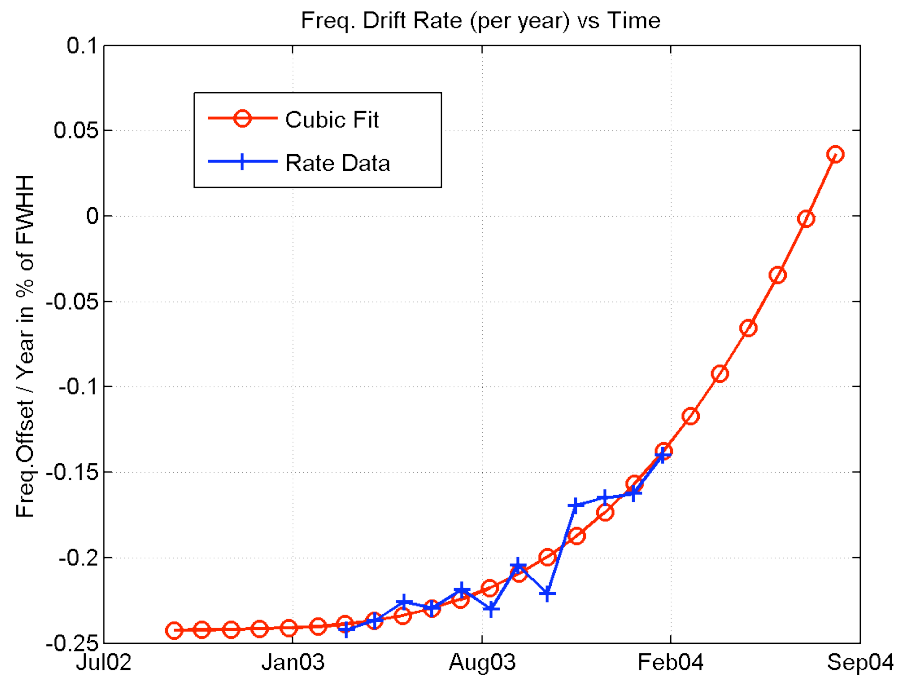
- Frequency calibration has 3 major terms:
  - Short term solar forcing: ascending/descending each with time variation that maps into latitude
  - Seasonal variation in above short term solar forcing, correlates with solar beta angle
  - Longer term drift
- We have performed a 2-year frequency calibration
  - Used UMBC's uniform clear L1b subset
  - Use sharp features in radiance due to  $\text{CO}_2$  and  $\text{H}_2\text{O}$ . (Avoid  $4.3 \mu\text{m}$   $\text{CO}_2$  band head.)
  - Compared  $B(T)$ 's computed from ECMWF to observed  $B(T)$ 's, shift frequency scale, via grating model, to minimize differences.
  - Bin monthly averages by latitude and day/night.
  - 7 arrays used to obtain average  $\Delta\nu$ .
  - M12 appears to be offset by  $1 \mu\text{m}$ .
- Matlab routine developed to correct frequency calibration errors
  - Uses computed radiances to determine local  $dB(T)/d\nu$  derivatives
  - Could be implemented as part of the RTA (Inputs: latitude, day/night, either month or solar beta angle, extrapolation of slow longer term drift).



Nov. 03 shift: 0.11% of width

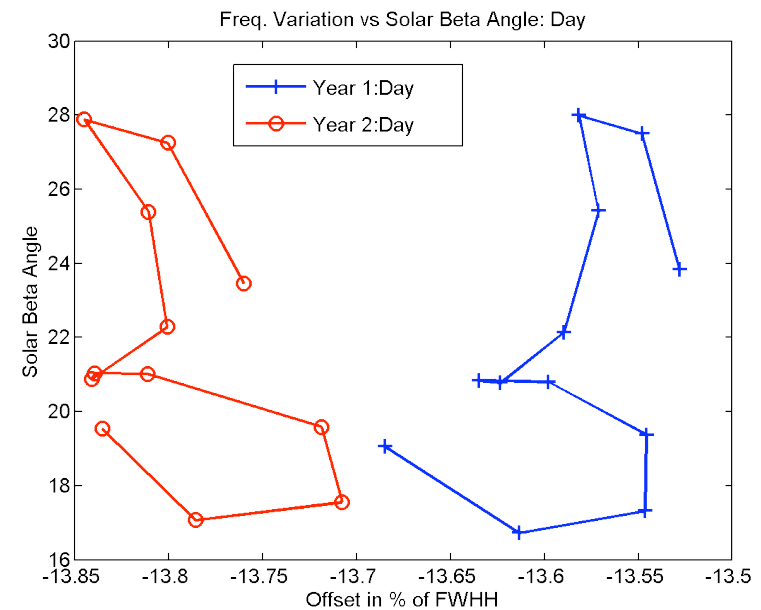
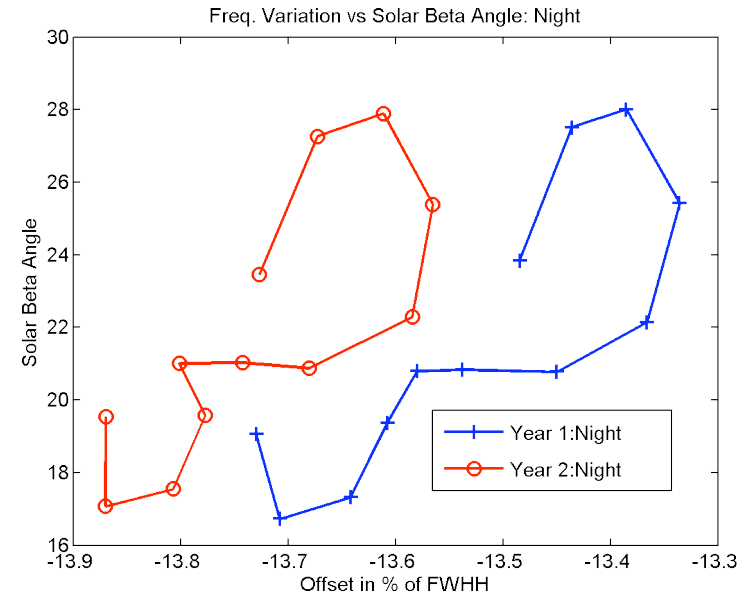
Day - Night  $\Delta v$  shows almost pure sinusoid

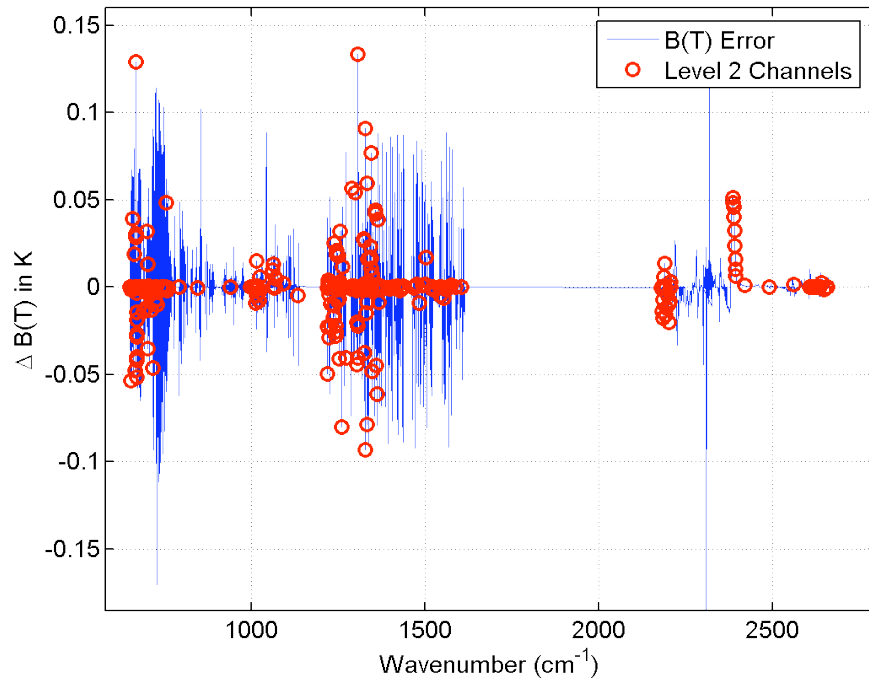
Although M12 is offset by 1% of a width from other arrays, it varies similarly in time



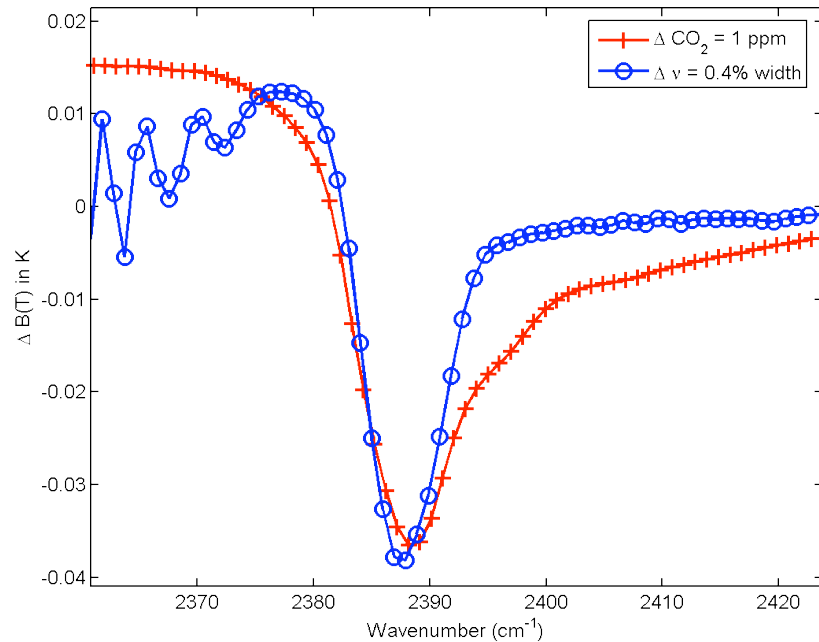
Total Freq. Variation:

0.3% orbital + 0.1% Nov. 03 + 0.8% slow drift ~ > 1% drift over life of mission.





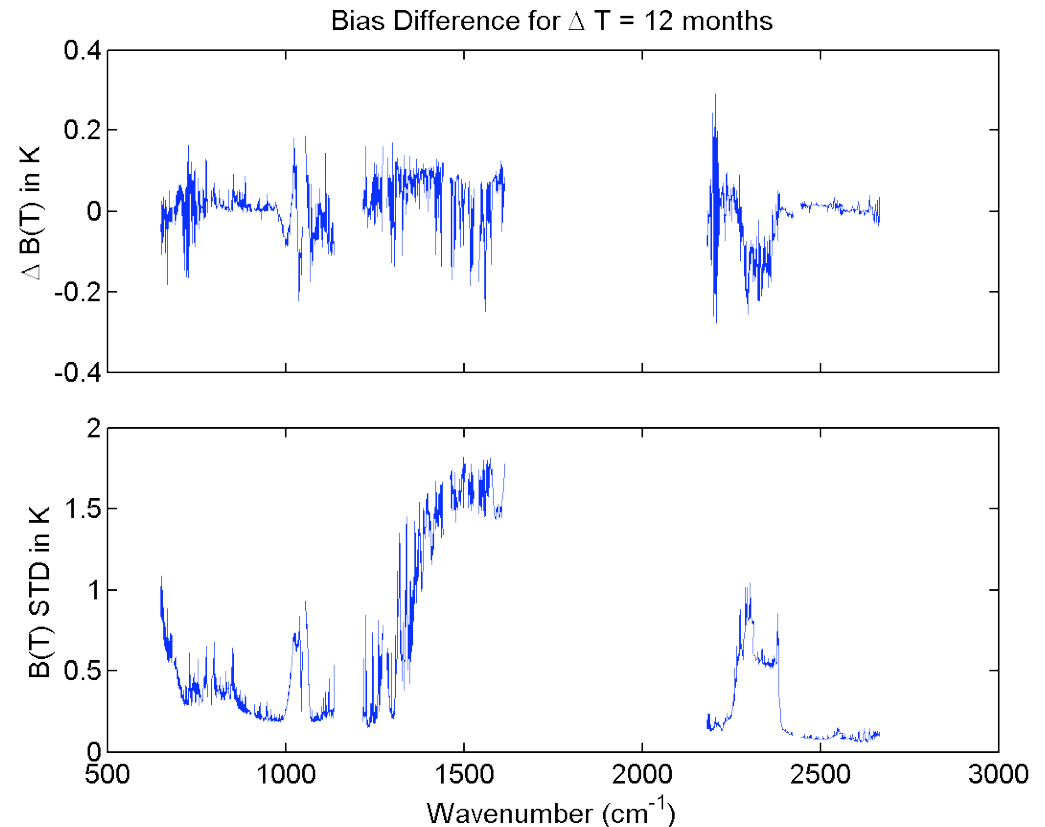
Note: Nov. 03 frequency shift of 0.11% width is easy to see in monthly mean biases relative to ECMWF for sensitive channels.



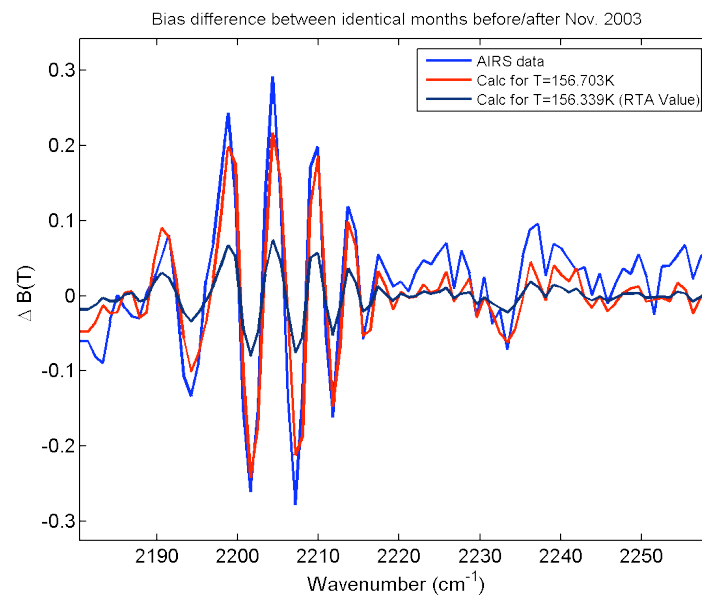
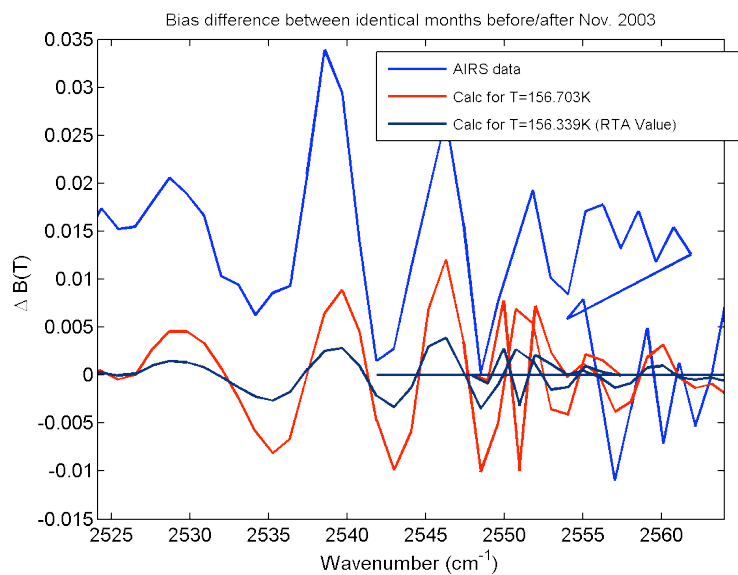
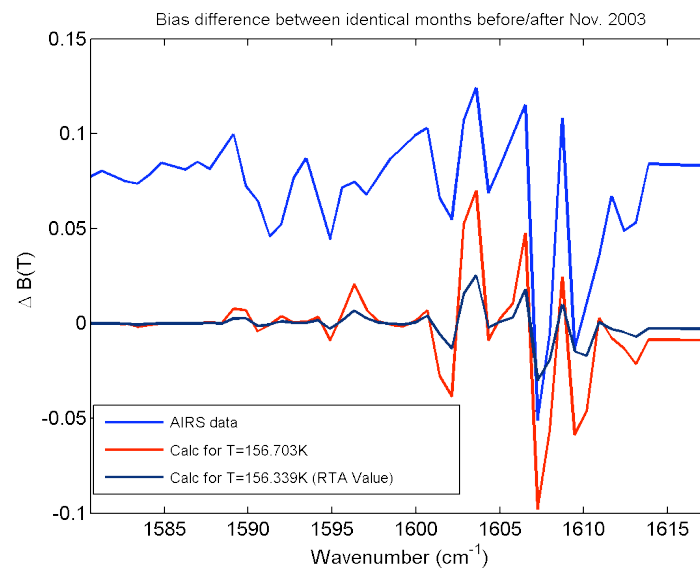
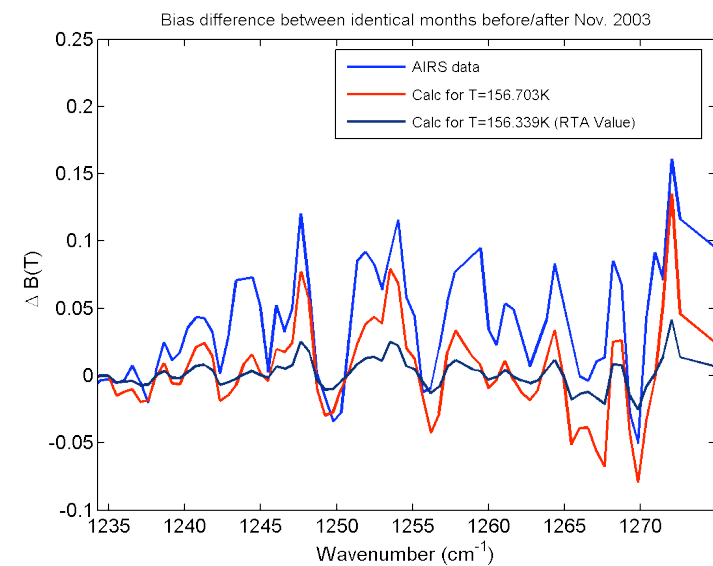
Difference between a frequency shift and variable CO<sub>2</sub> almost impossible to separate. Note that the 4.3 μm channels are very good for CO<sub>2</sub> due to low water sensitivity.

# Fringes

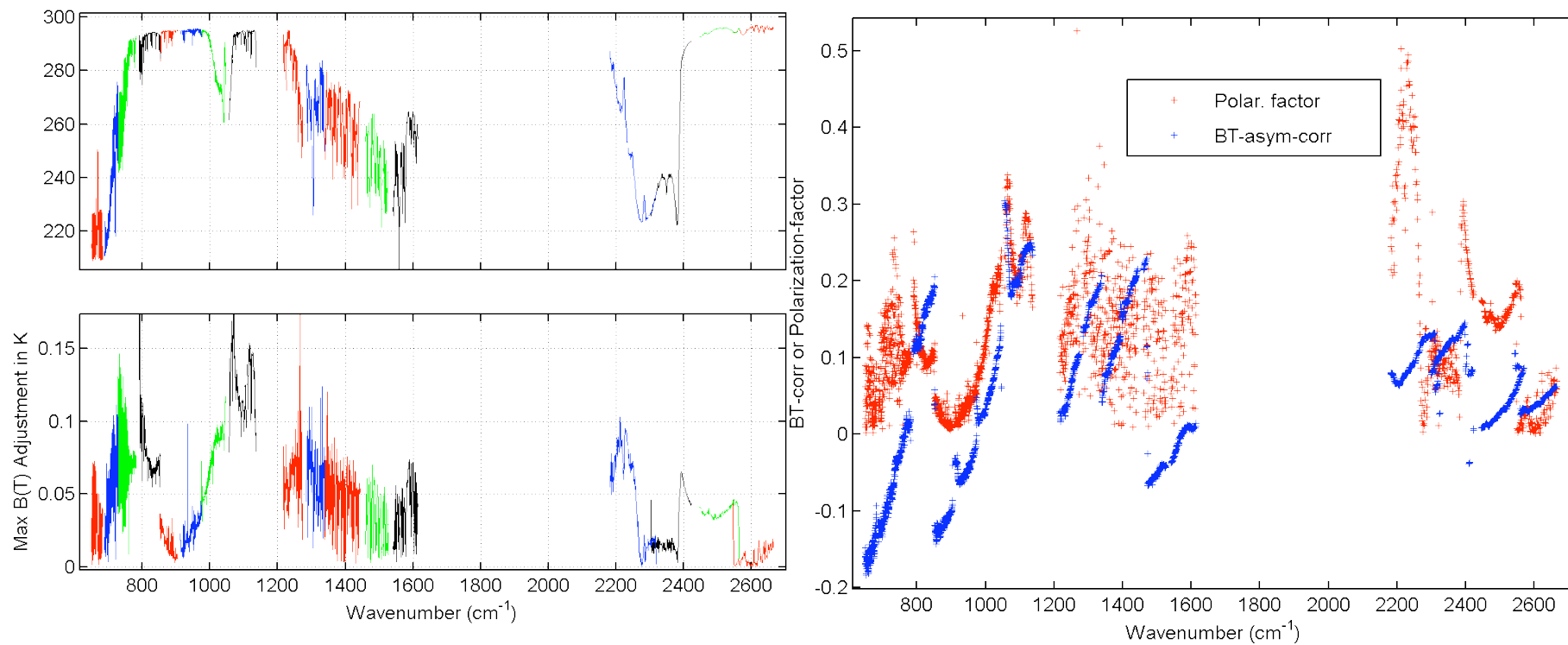
- Fringes moved due to Nov 03 shutdown
  - Goal was to keep frequencies unshifted
  - Resulted in different temperature for filter producing fringes
- Somehow, we got the wrong filter temperature when producing the post Nov. 03 RTA
- Moreover, the decision was made to have only one RTA for reprocessing, using the supposed post-Nov. 03 fringe positions
- So fringes are incorrect for all AIRS data



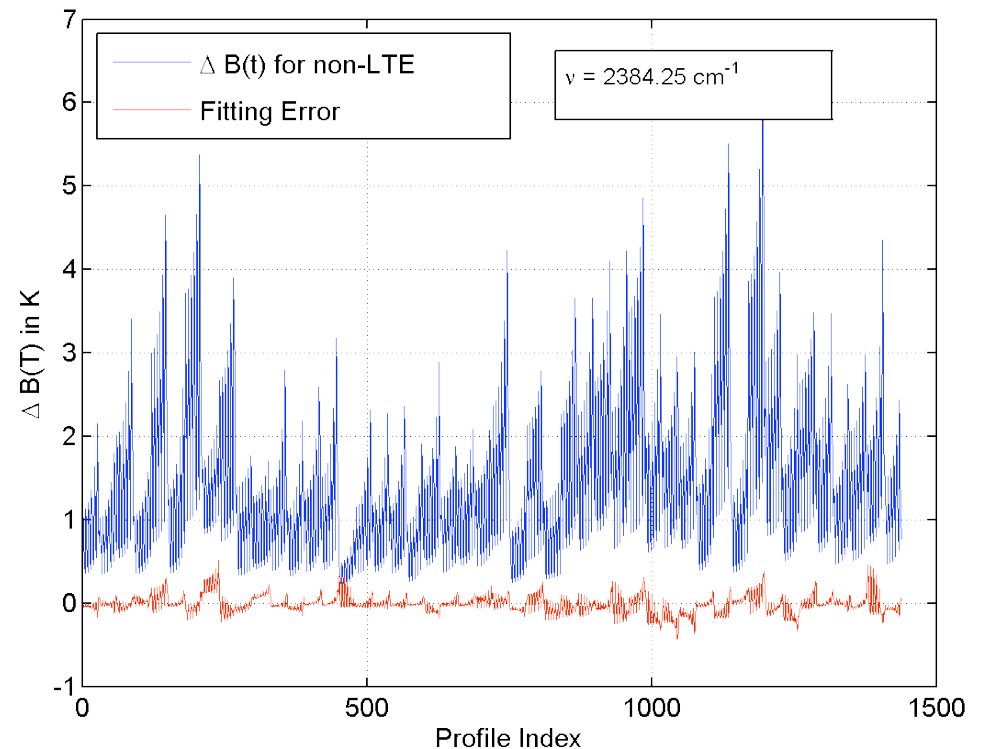




# Max Scan Bias Asymmetries



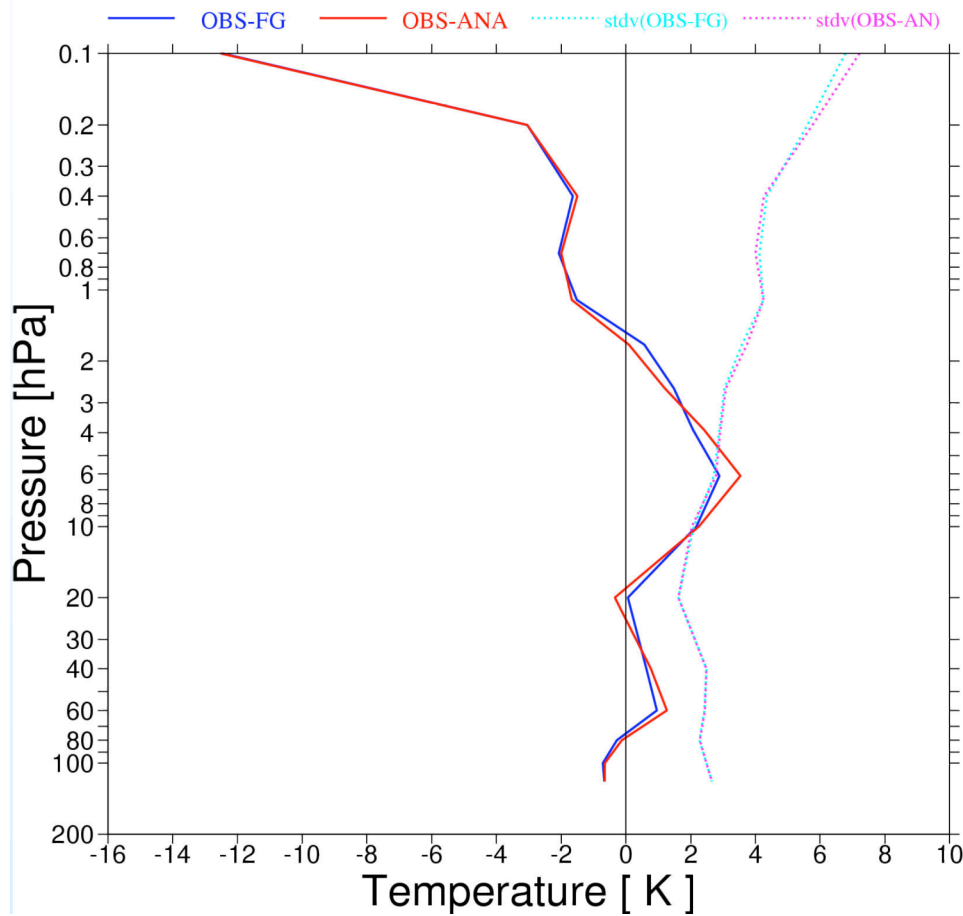
- Some work on fast non-LTE model.
- Fast parameterization looks good, fundamental theory being tested
- First principles calculations are relatively good, but need non-LTE vib/rot temperatures, not a simple calculation
- Non-LTE small for  $\sim 2380 \text{ cm}^{-1}$  region: corrections should be easy
- Various possibilities:
  - Use 15 micron channels in regression for 4 micron non-LTE along with solar angle
  - Use strong non-LTE in  $2330 \text{ cm}^{-1}$  region to predict non-LTE in  $\sim 2380 \text{ cm}^{-1}$  region (use ECMWF to estimate amount of non-LTE near  $2330 \text{ cm}^{-1}$ ).
- Does anyone care?



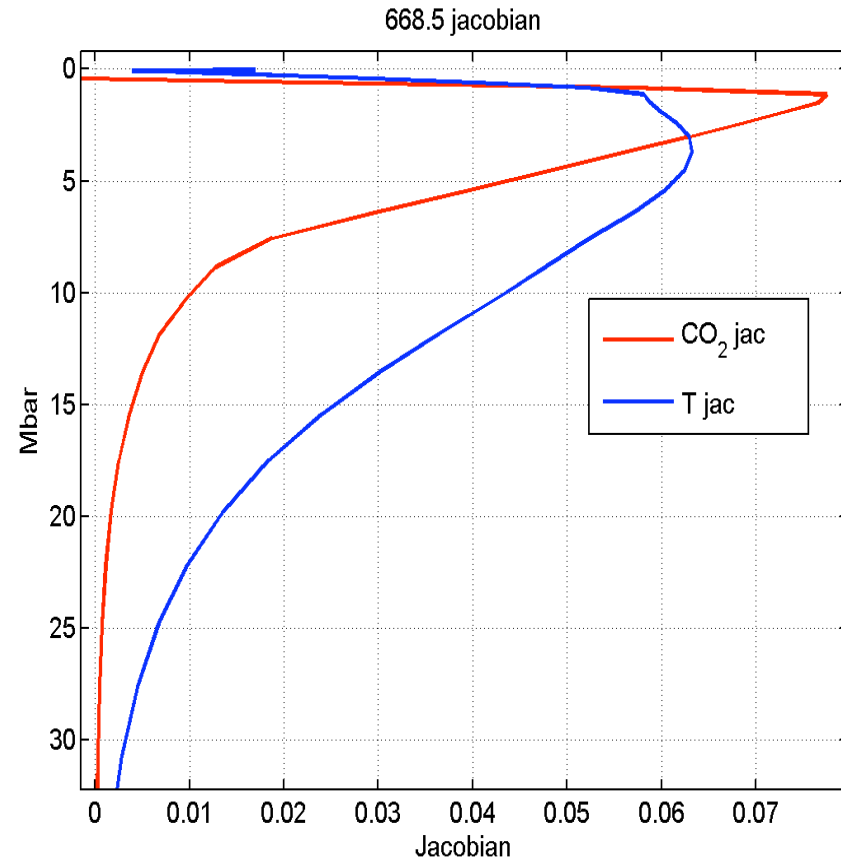
## Variable Gases

- $\text{CH}_4$  and  $\text{N}_2\text{O}$  can vary significantly, including in the stratosphere where AIRS channels have sensitivity.  $\text{CO}_2$  can vary slightly as well.
- I have observed many variations in  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ , and  $\text{CO}_2$  channel biases (vs ECMWF) with latitude. These biases generally vary with the channels stratospheric sensitivity.
- Sensitivity studies using MIPAS constituent retrievals for  $\text{CH}_4$  and  $\text{N}_2\text{O}$  show some significant AIRS sensitivities.
- Are highest altitude channel biases dominated by ECMWF (esp. for  $\text{CO}_2$ )?
- Biases change character as go to lower peaking channels
  - Due to variable  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ ,  $\text{CO}_2$ , often in stratosphere?
  - $\text{CO}_2$  from 791.7 and 2390  $\text{cm}^{-1}$  show excellent agreement with CMDL, including almost perfect variation with season at 50 degrees latitude.
- Much work needs to be done. Hope to utilize MIPAS monthly mean profiles for validation.
- These effects could pollute latitudinal dependence of AIRS products.

# MIPAS for High Altitude RTA Validation?



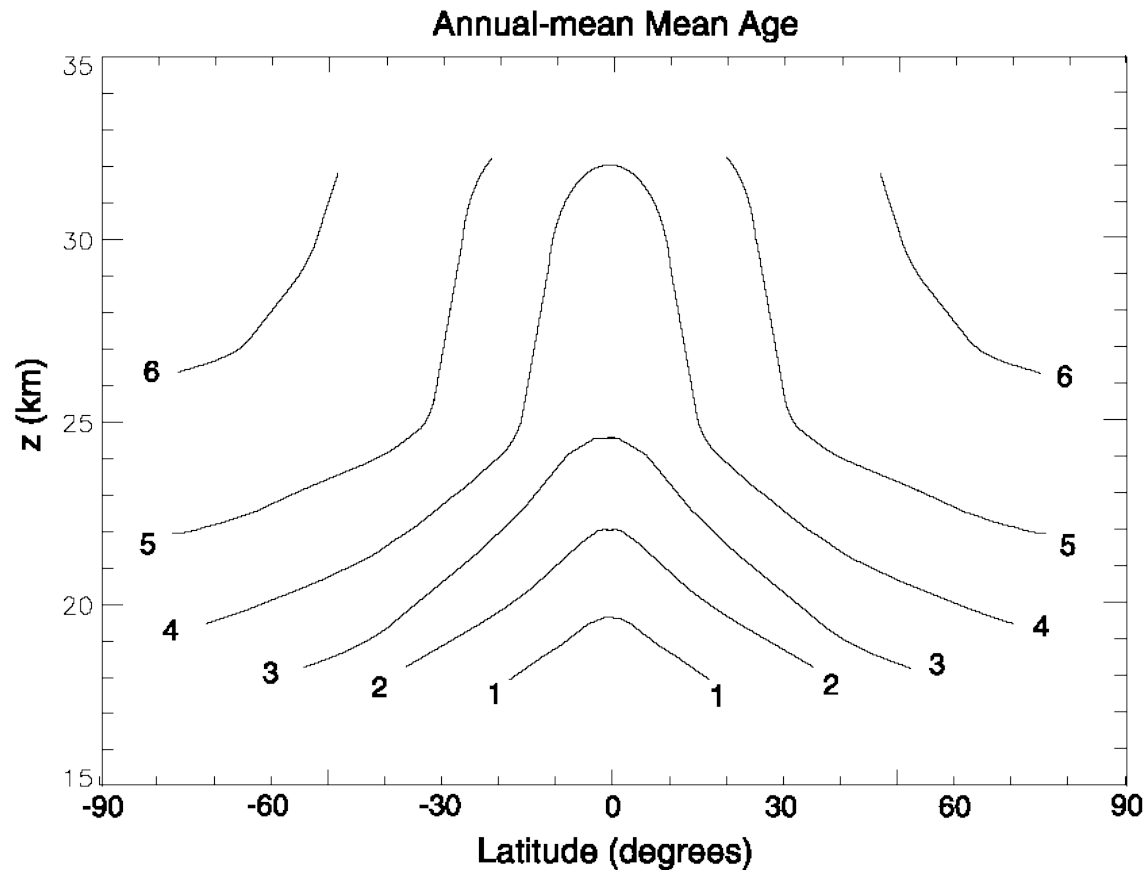
MIPAS - ECMWF



Hopefully can get global monthly mean profiles from MIPAS (Oxford) for T, CH<sub>4</sub>, N<sub>2</sub>O

# Stratospheric Variability

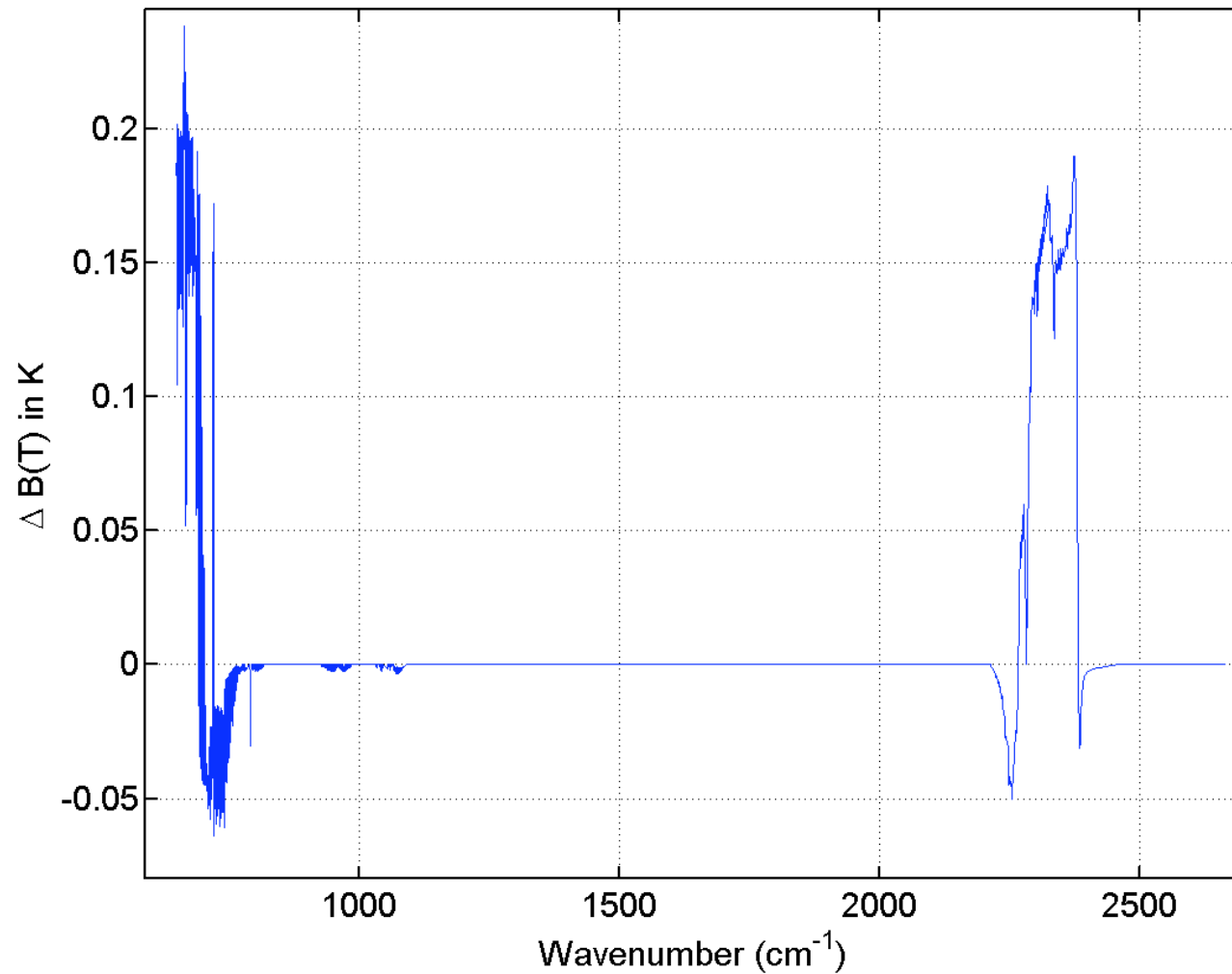
Due to long-time scale of trop to strat exchange



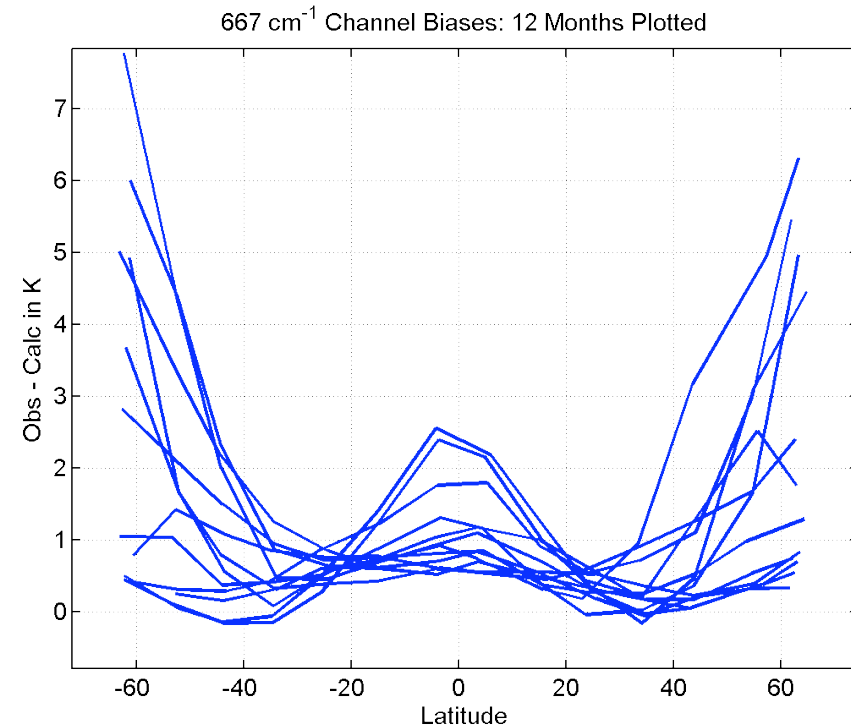
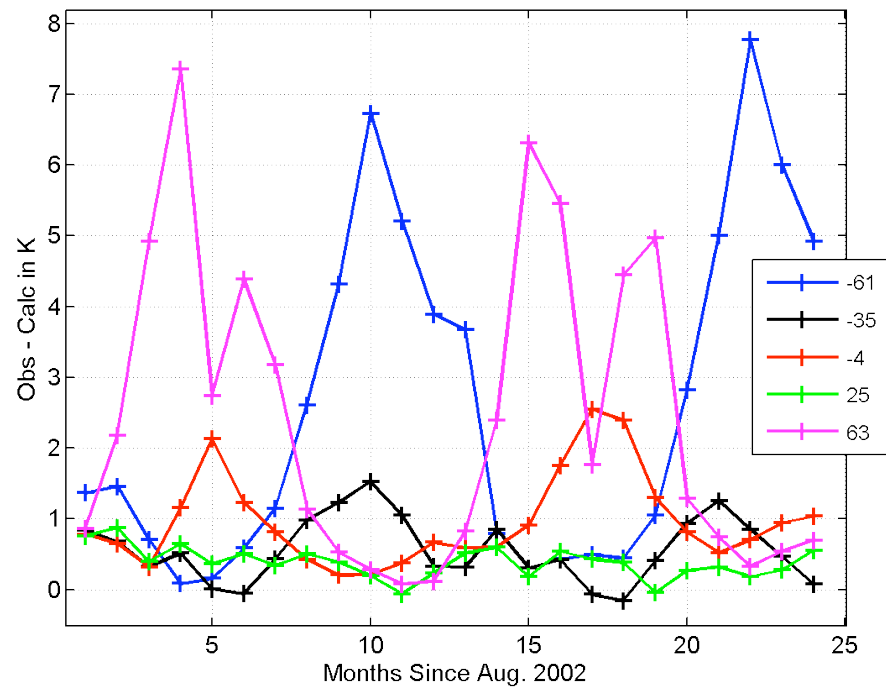
**Figure 7.** Schematic diagram of the altitude-latitude distribution of the annually averaged zonal mean of mean age based on vertical profiles from observations listed in Table 1 and ER-2 measurements around 20 km (e.g., Figure 6a). As it is based almost exclusively on Northern Hemisphere data, the schematic is hemispherically symmetric.

## Rough Estimate of Stratospheric $\text{CO}_2$ Sensitivity

$\text{CO}_2$  varied in stratosphere using  
nominal ER-2 measurements



## Observed 667 cm<sup>-1</sup> Biases versus ECMWF



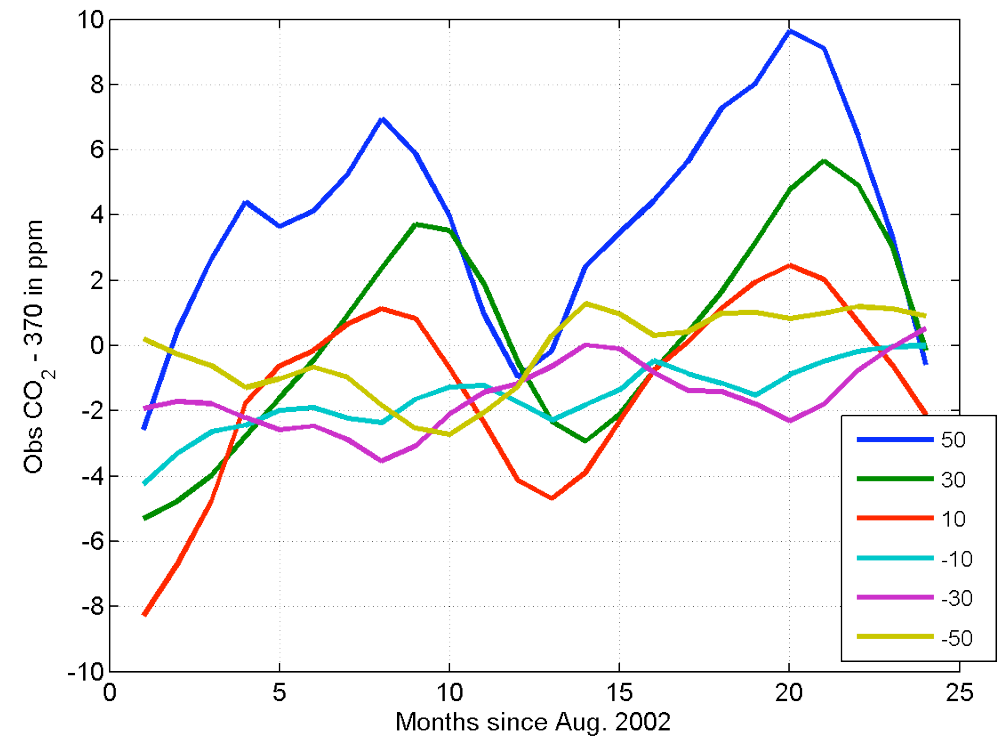
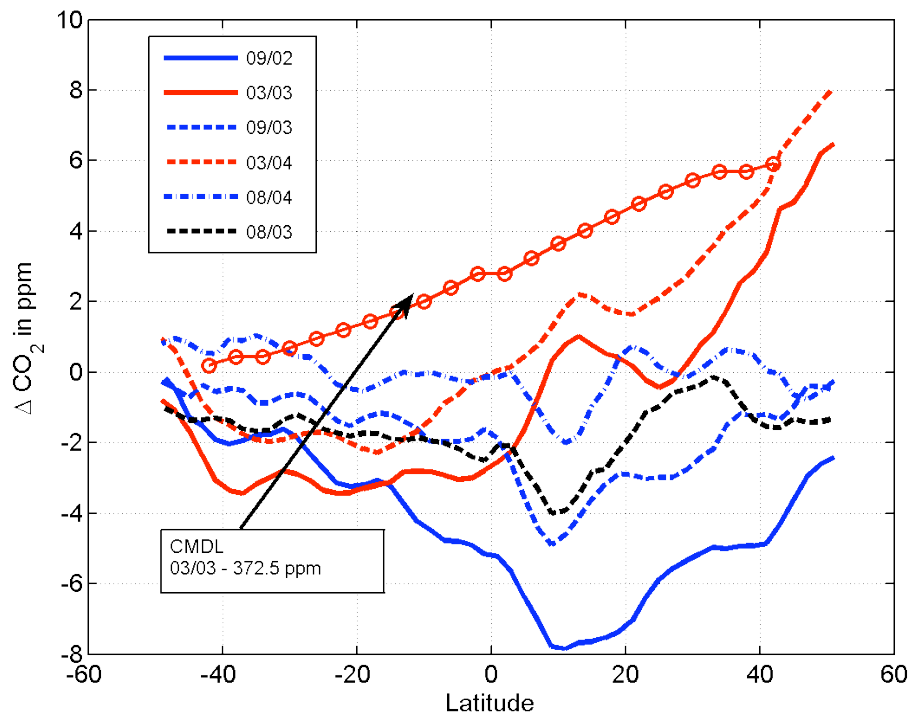
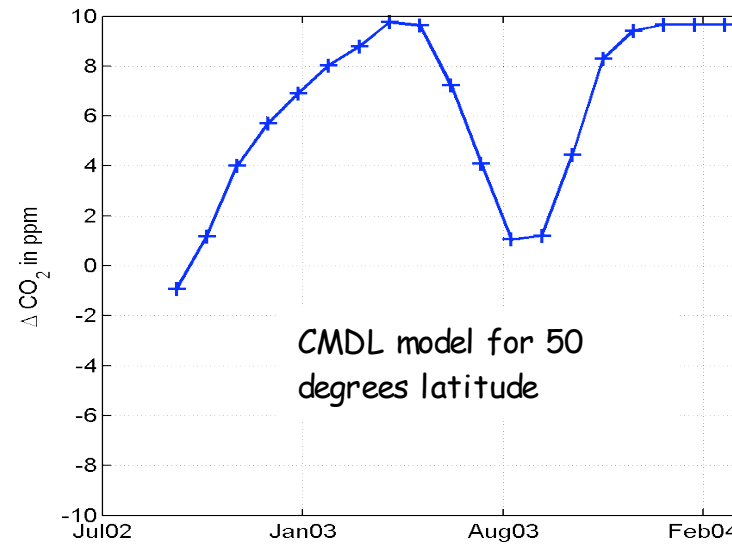
- Biases much larger than expected for stratospheric CO<sub>2</sub> variability, esp at poles
- Phase reversal between poles with time. 25 deg N very stable bias
- Unsure if biases are too large (2K) relative to MIPAS, need more details on latitude range of MIPAS biases relative to ECMWF
- Any suggestions?

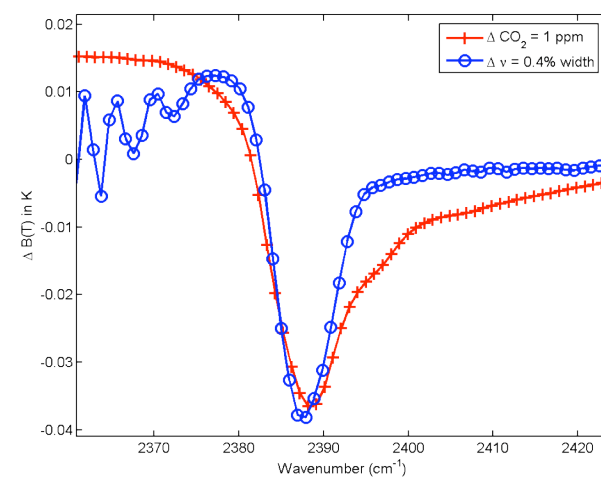
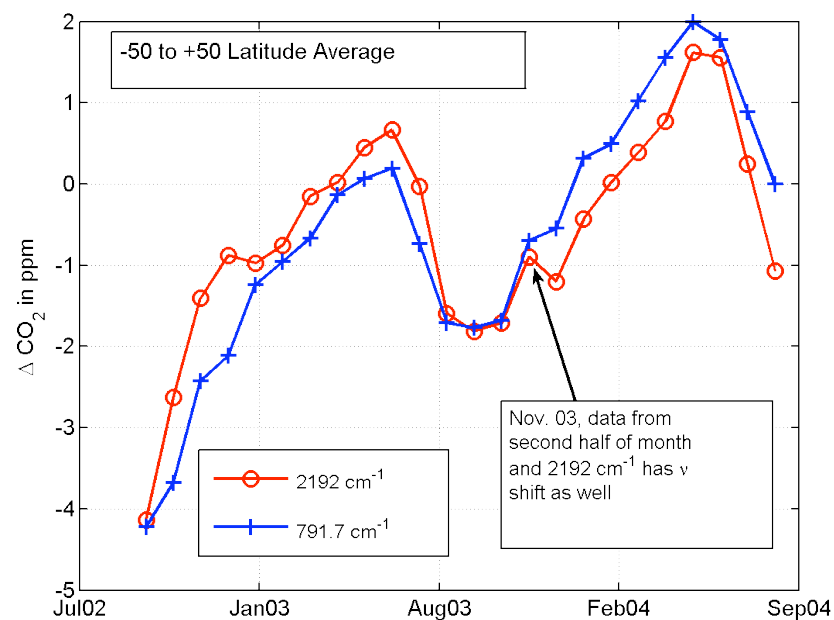
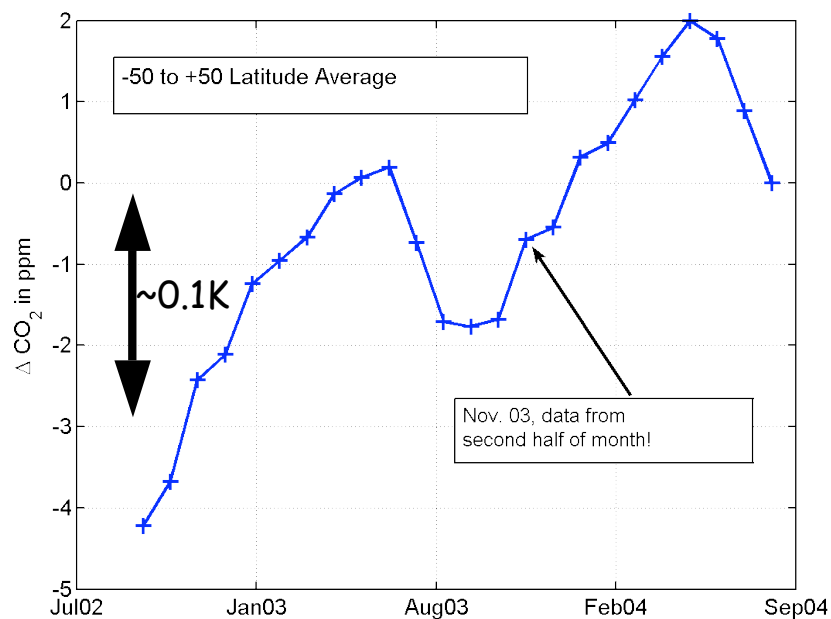


Data reproduces basic form of  
CMDL models

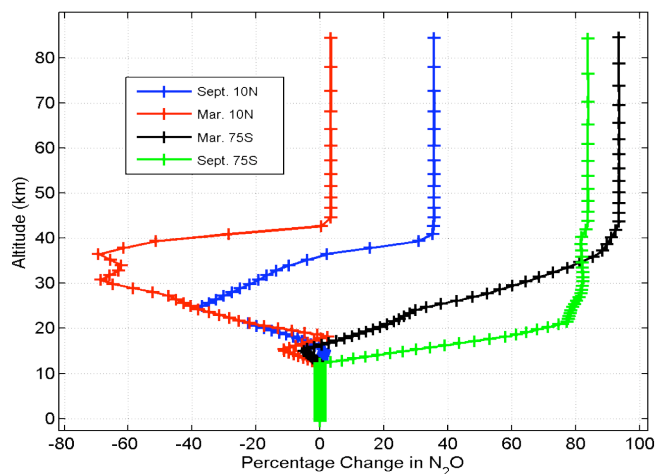
Need a single overall calibration,  
but within ~2-3 ppm initially

Is fine structure real?

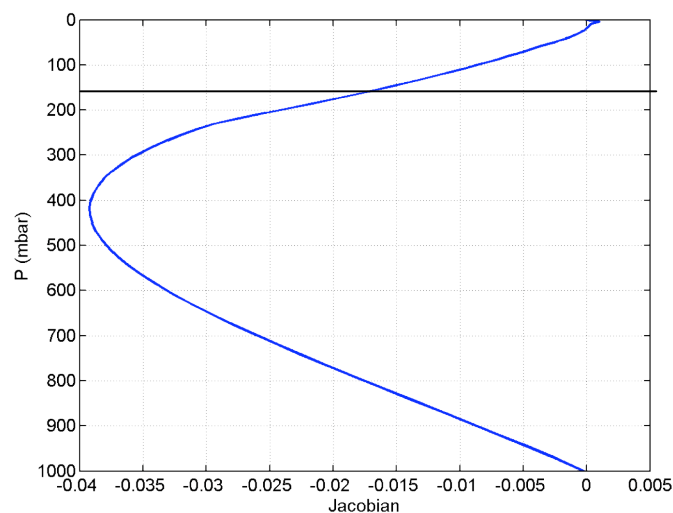




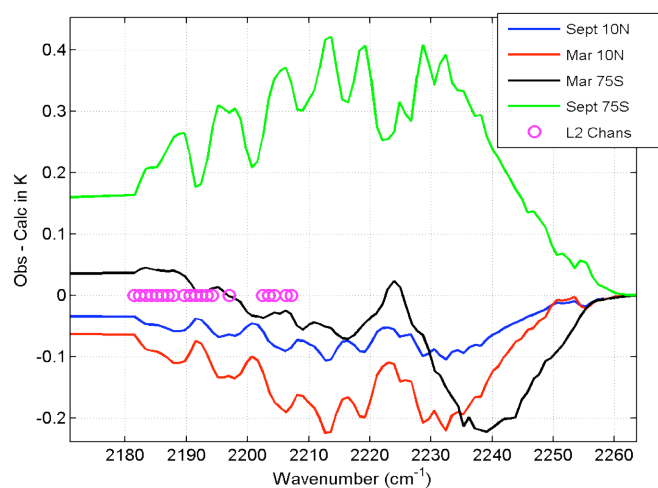
Nominal MIPAS Variability



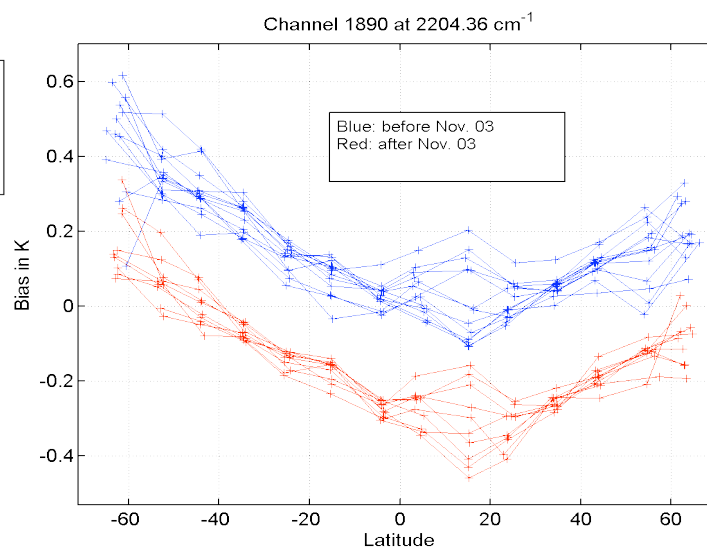
N<sub>2</sub>O Jacobian for 2204 cm<sup>-1</sup>



Computed Bias from Above Profiles

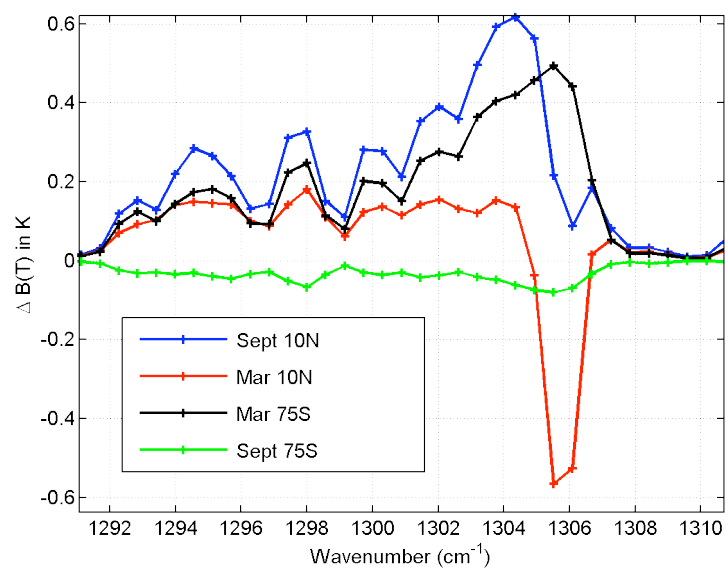
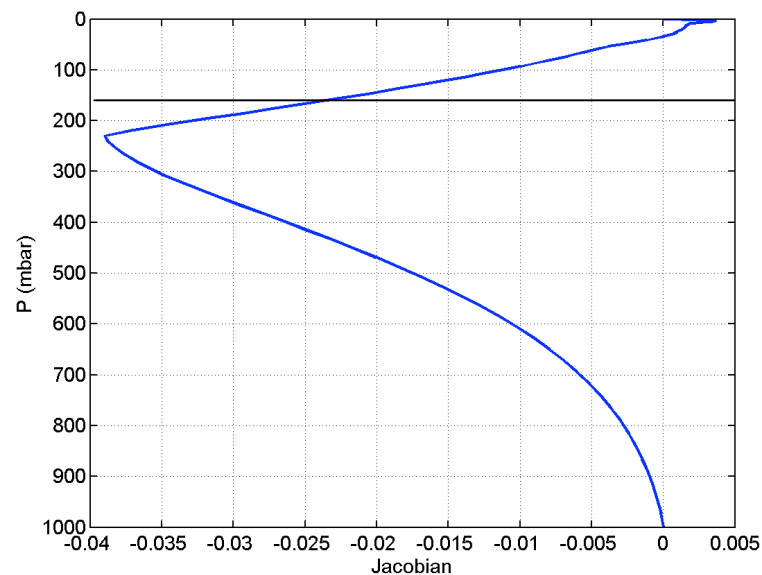
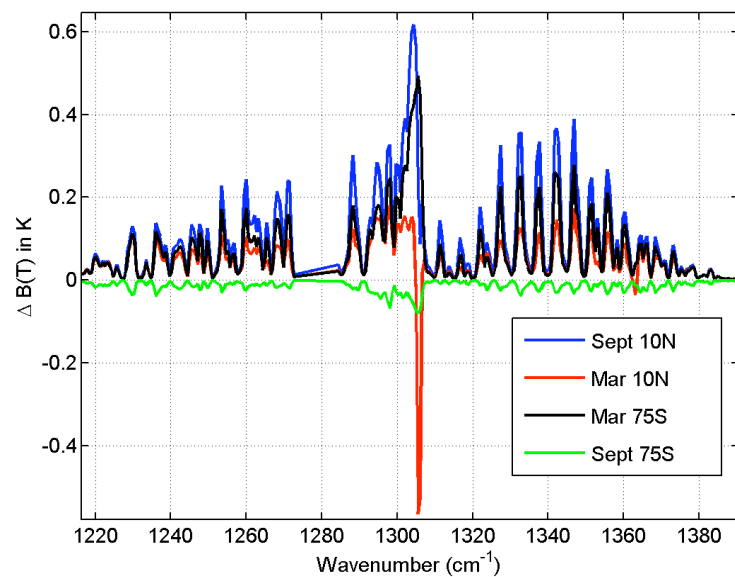


Observed Biases vs ECMWF



- Need to let N<sub>2</sub>O vary in RTA?
- Sounding channels impacted by variable N<sub>2</sub>O?
- Fringing effects long-term signal

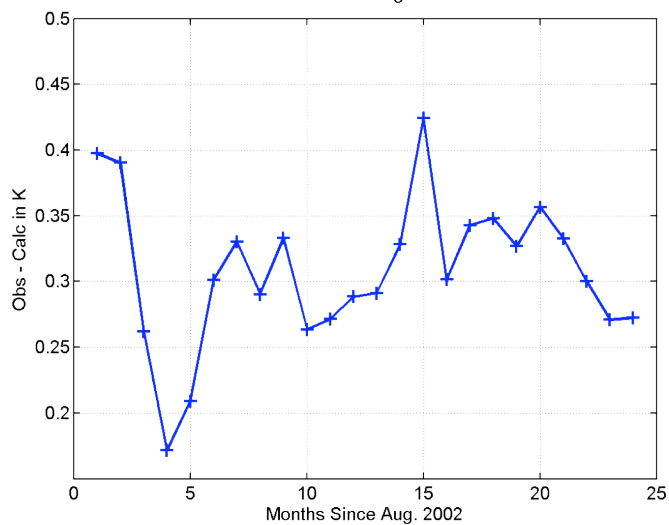
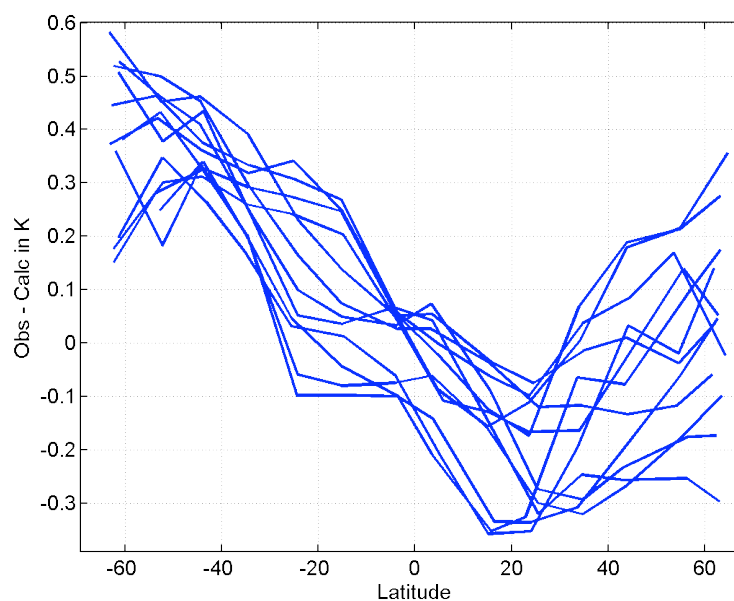
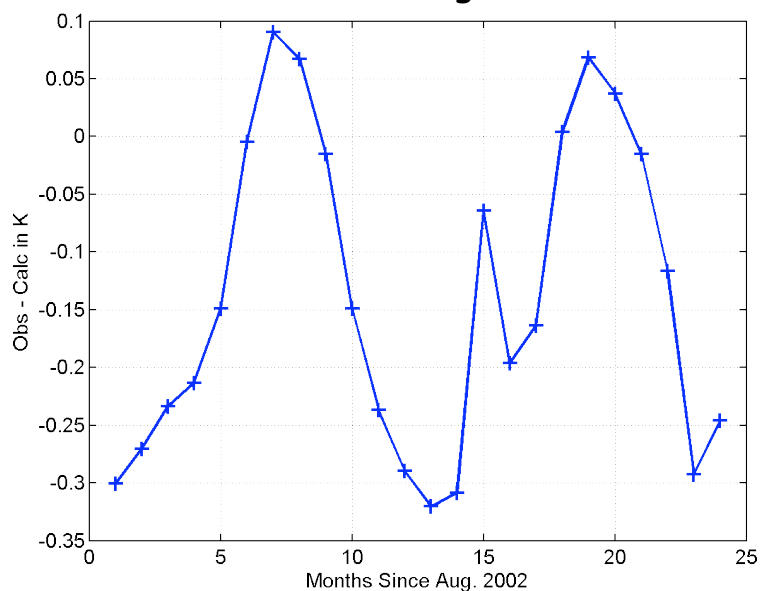
# CH<sub>4</sub> Stratospheric Variability (from MIPAS)



# Observed Biases versus ECMWF in CH<sub>4</sub> Channels

CH<sub>4</sub> channel at 1304 cm<sup>-1</sup>

Bias for 34 deg. N.

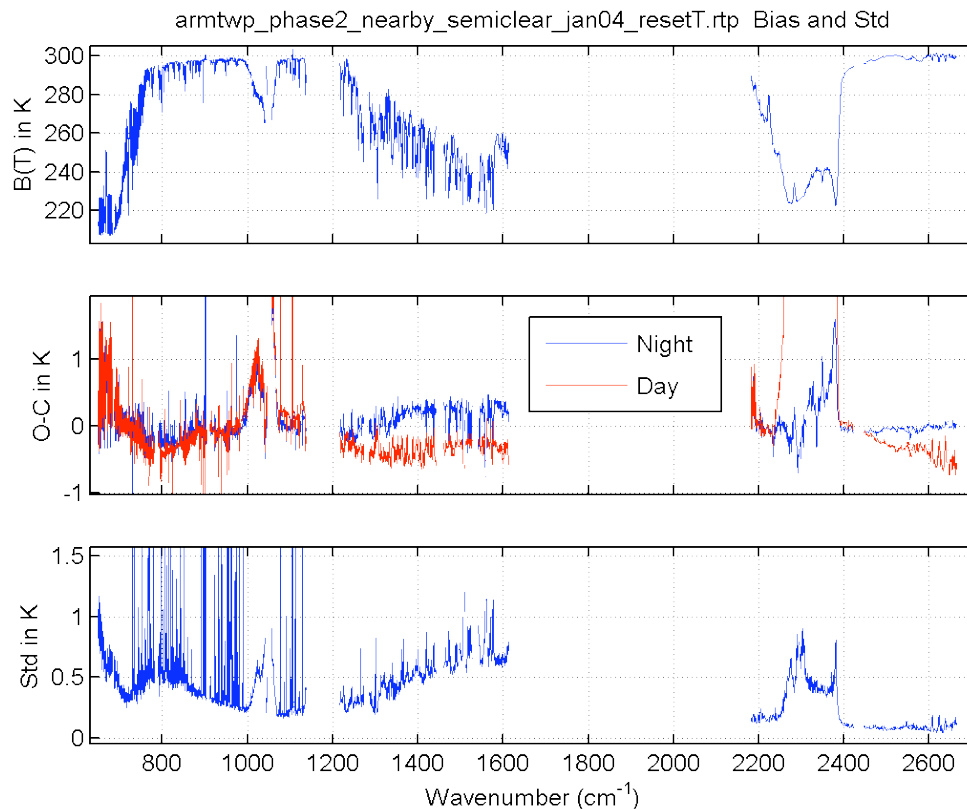


717 cm<sup>-1</sup> channel bias for 34 deg. N.  
This channel's CO<sub>2</sub> Jacobian is very  
similar to the 1304 cm<sup>-1</sup> channel CH<sub>4</sub>  
Jacobian.

New results: Multiple phases helps with error analysis, priority for special issue

RS-90 sondes

Sonde calibration continuing - no Milosevich corrections used here



Clear determination from H. Aumann global SST studies. Probably lets through  $\sim 0.3\text{K}$  cloud signal, on average.

Fit for SST (minimizes clouds)

RTA from ARM-TWP 2002/2003

Used "global" clear-flag

$\sim 5\%$  FOVs survived clear test

30-50 mm  $\text{H}_2\text{O}$ , = 7-11K depression at  $800\text{ cm}^{-1}$

ECMWF for T above sondes

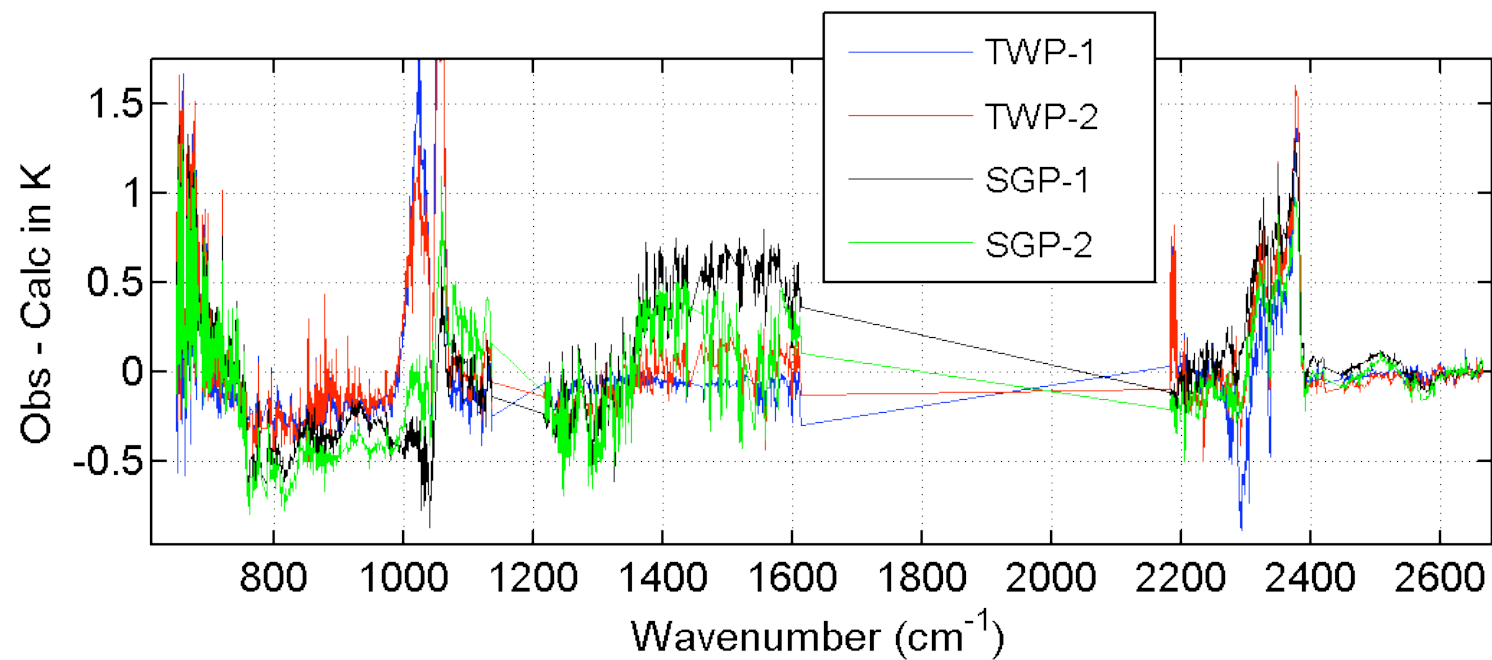
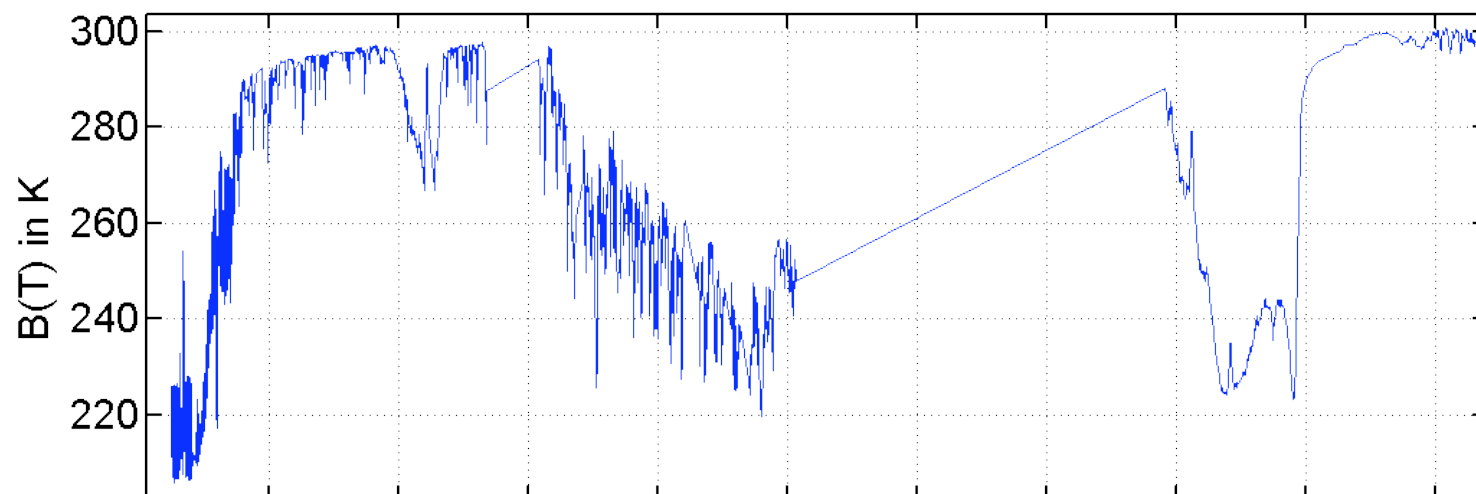
Most clear from Fall 2003

Brand new data, so preliminary, but probably best estimate of RTA error bounds.

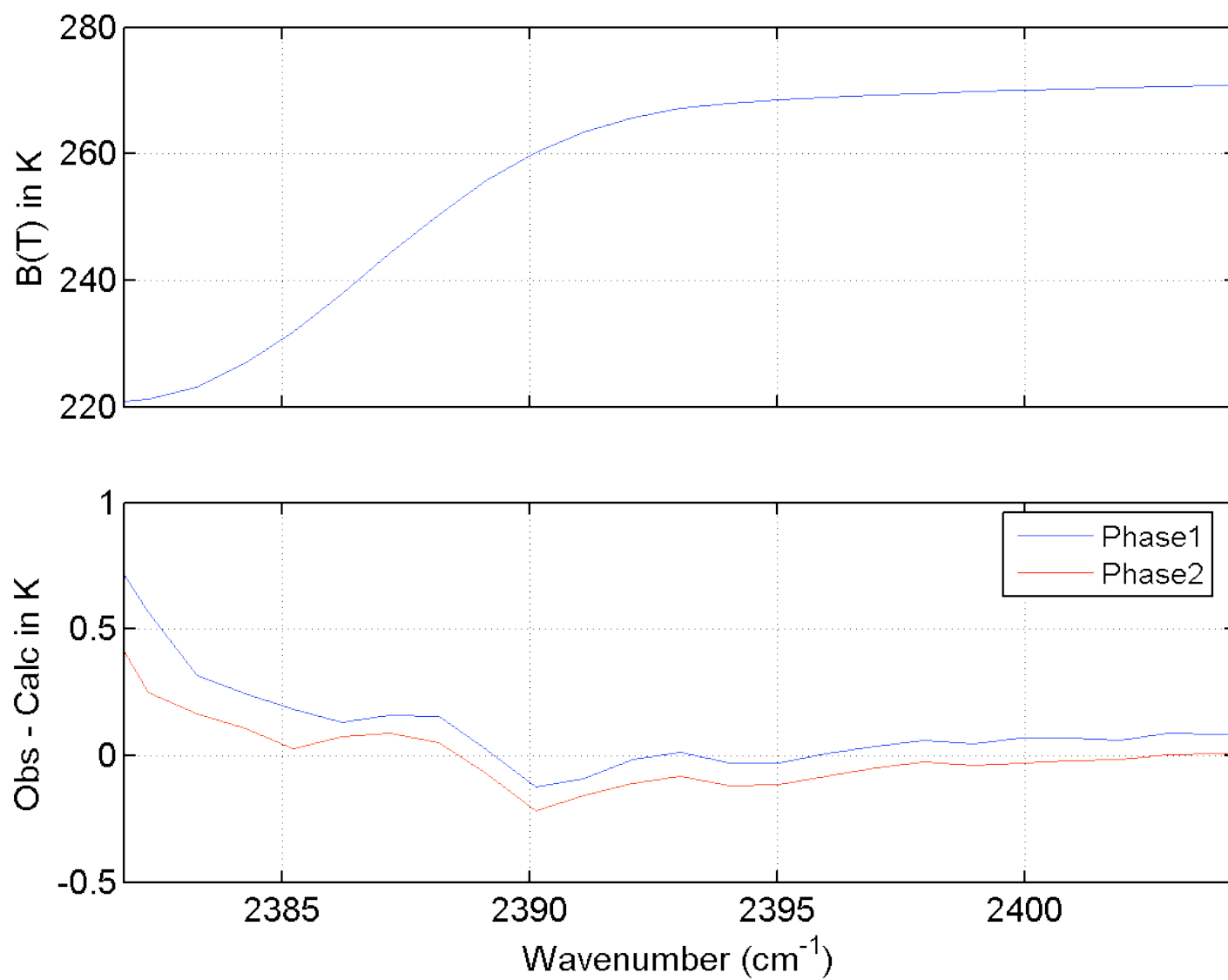
$800\text{-}1000\text{ cm}^{-1}$  Errors =  $\sim 2\%$  water

$1400\text{ - }1600\text{ cm}^{-1}$  Errors =  $\sim 3\%$  water

# ARM: 2 of 3 Phases Ready

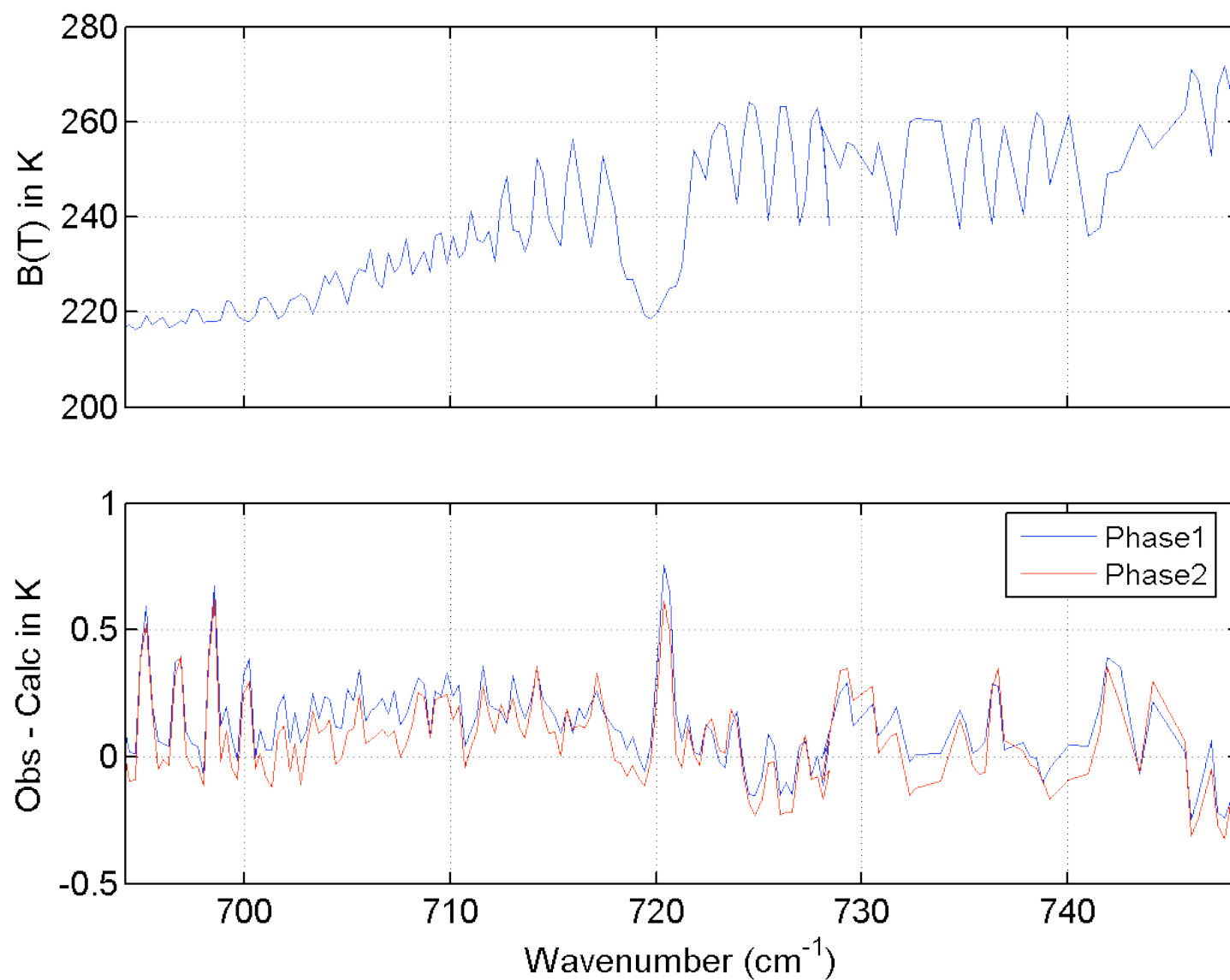


# SGP Short Wave

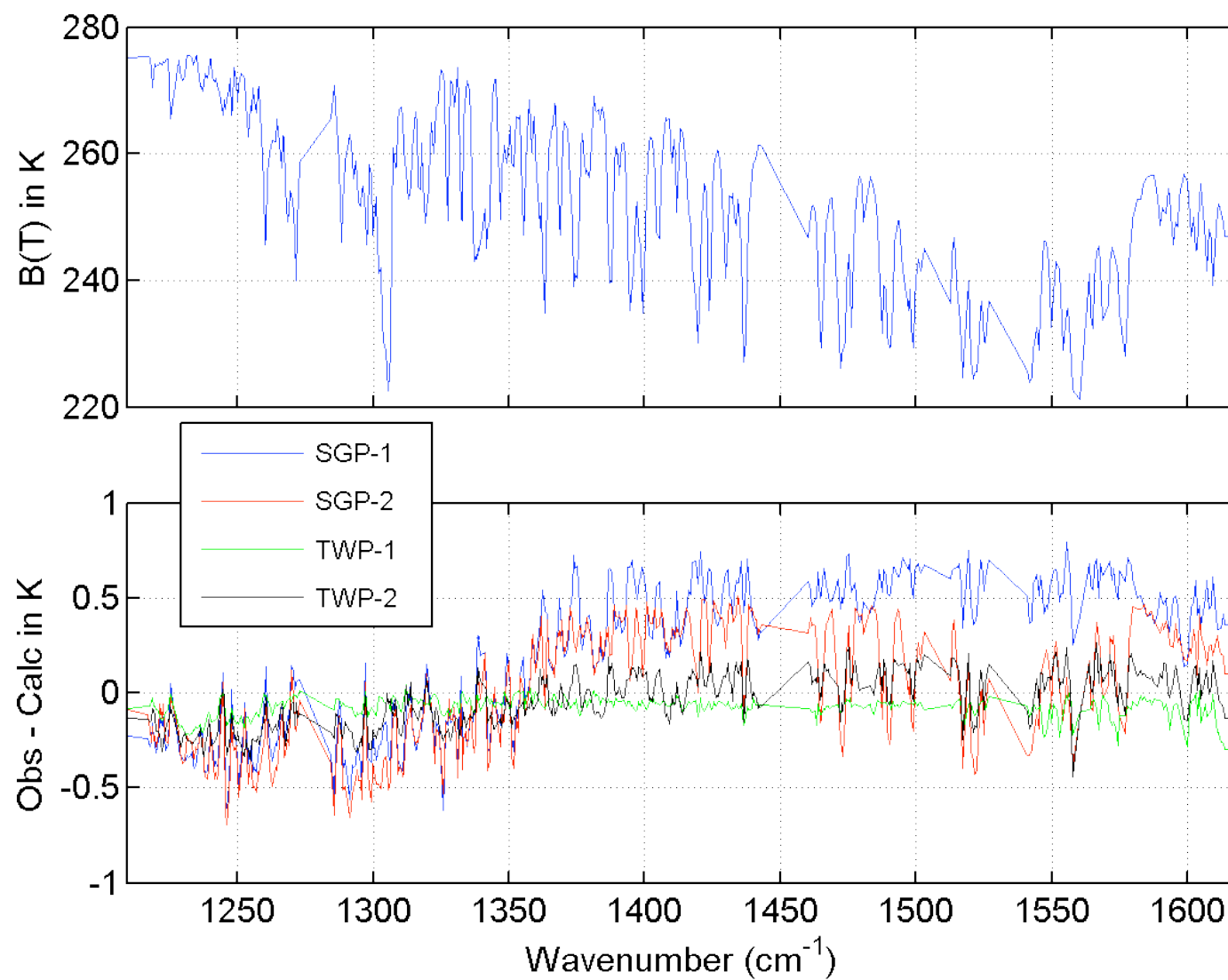




# SGP Long Wave

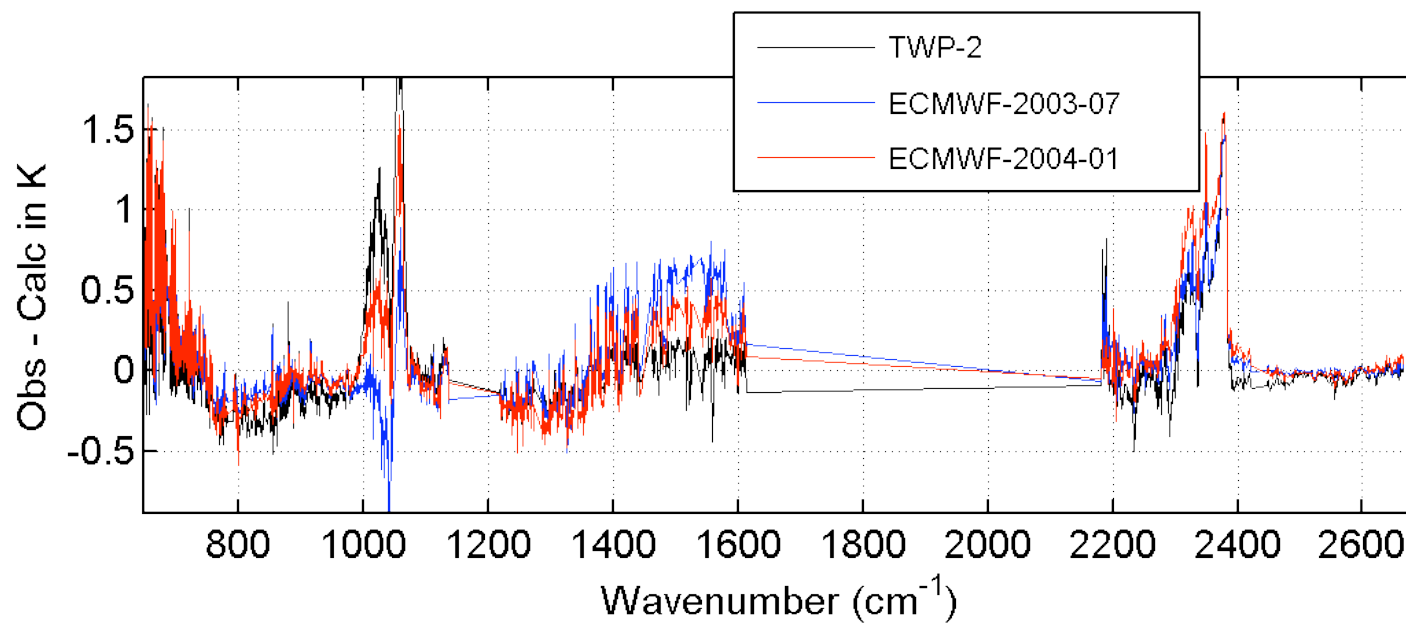
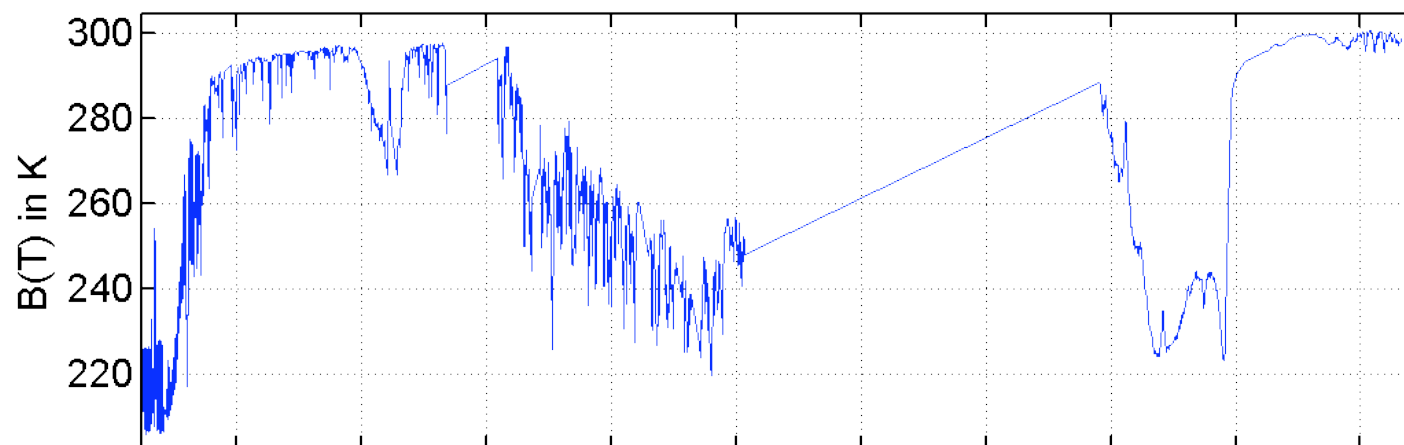


# SGP and TWP Water Region

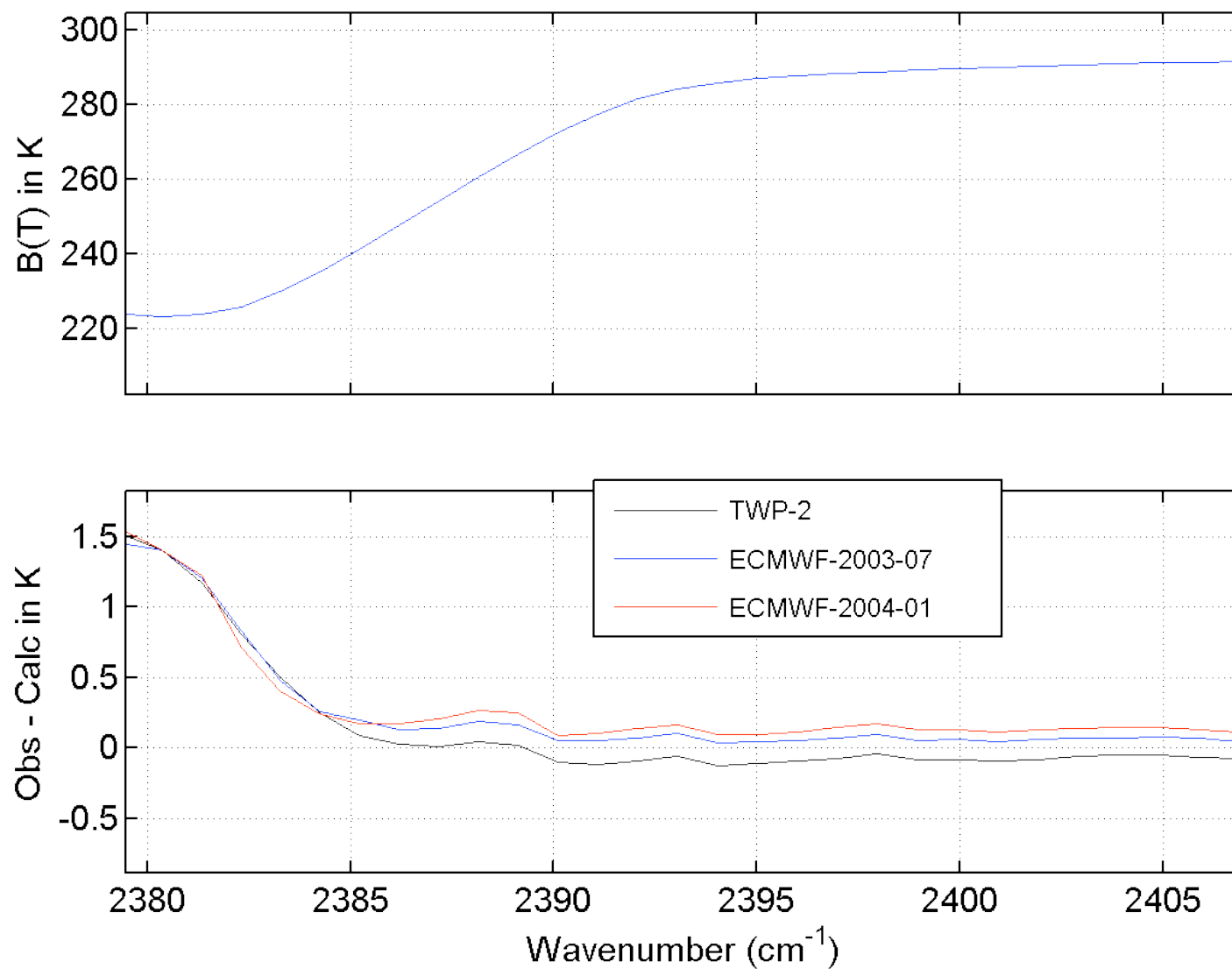


# TWP versus ECMWF

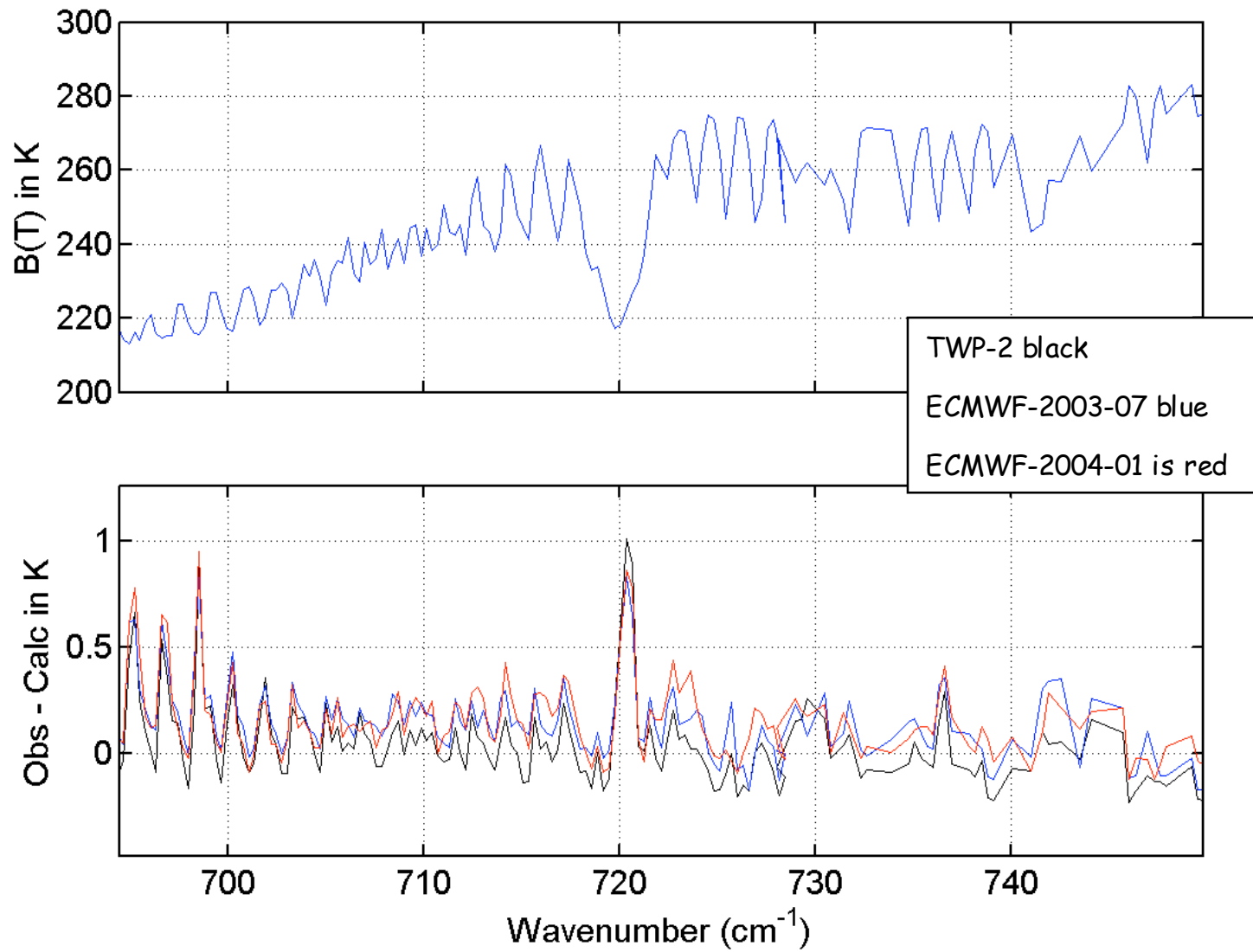
(ECMWF averaged over ~10-40 deg. Latitude)



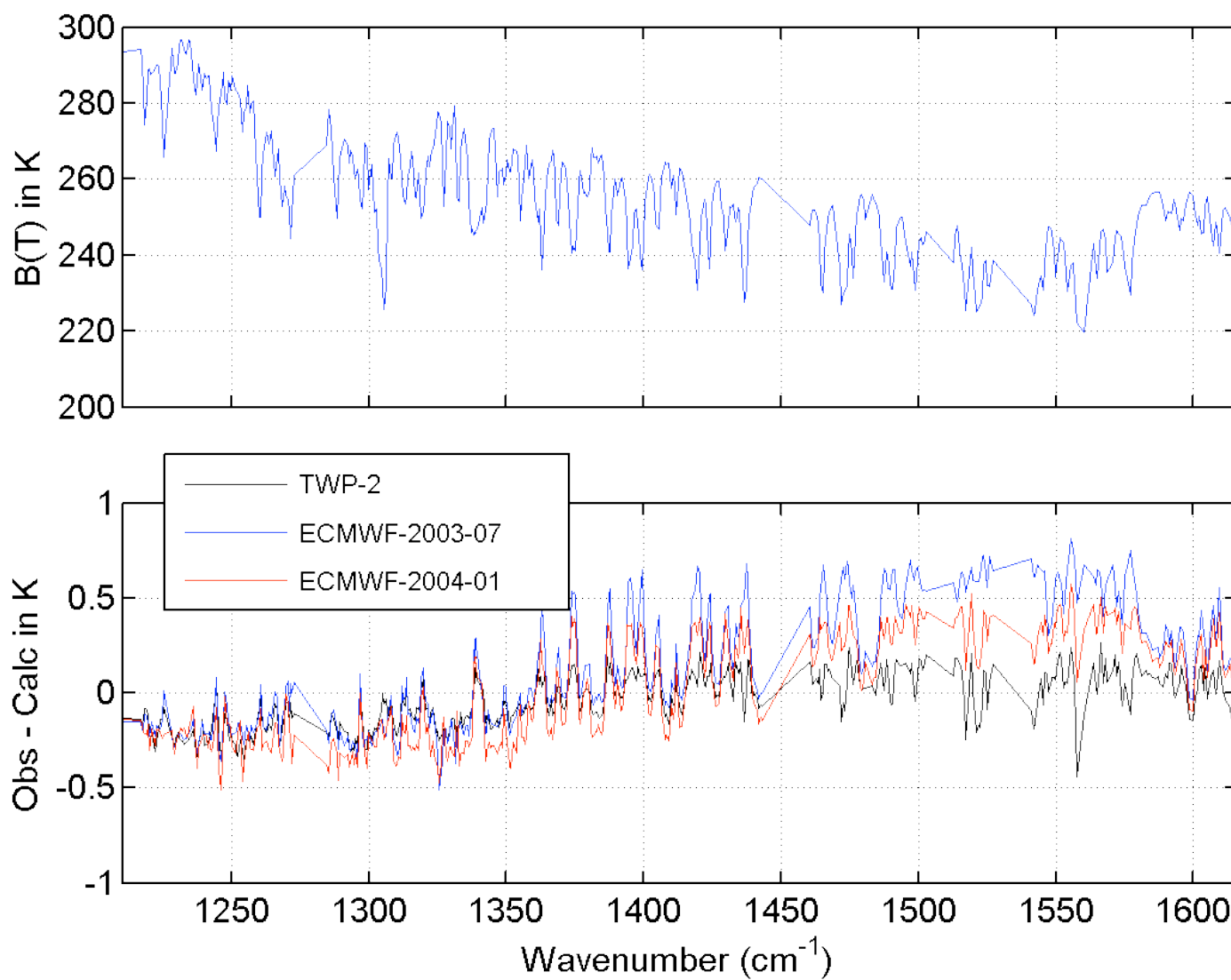
# TWP versus ECMWF



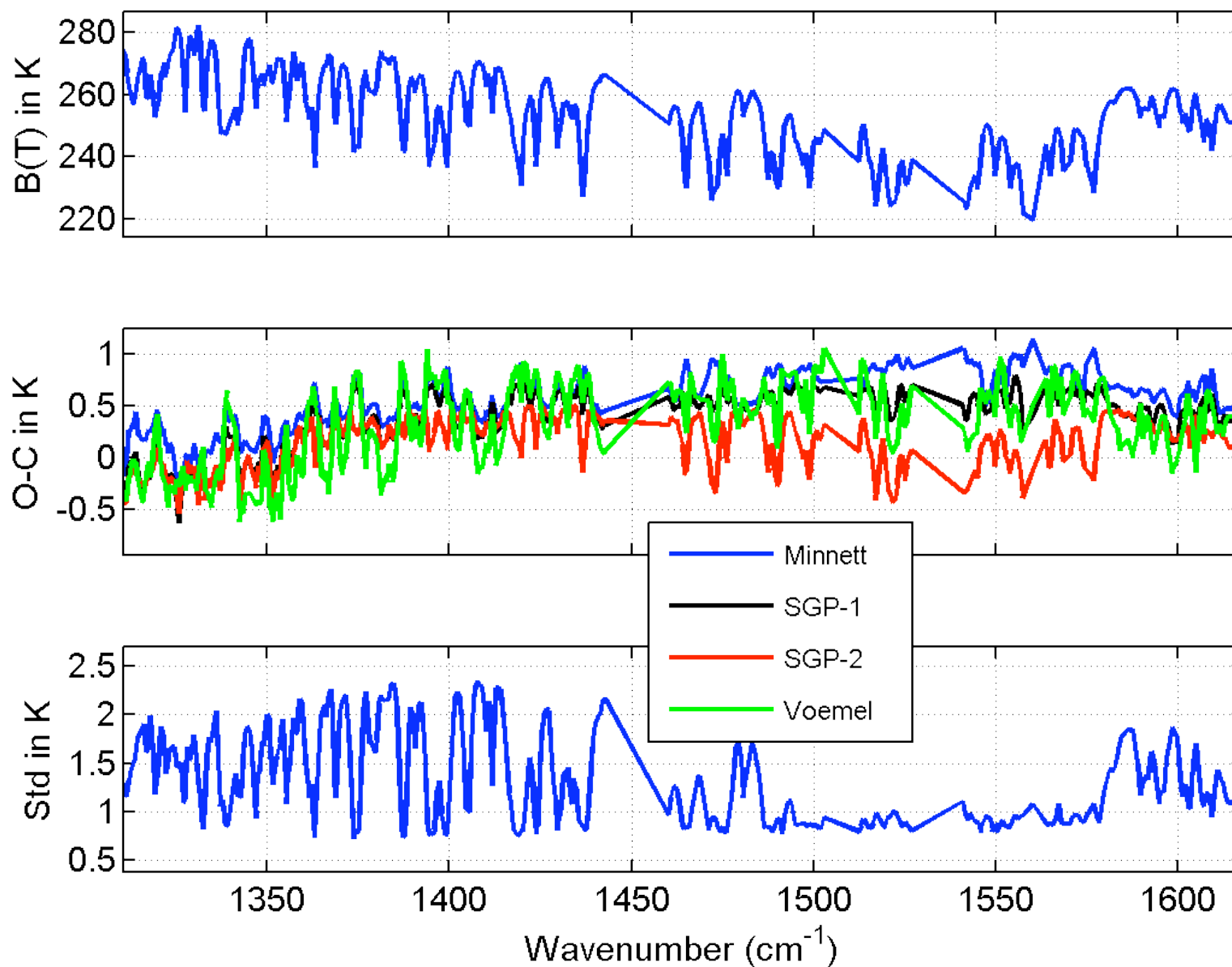
# TWP versus ECMWF



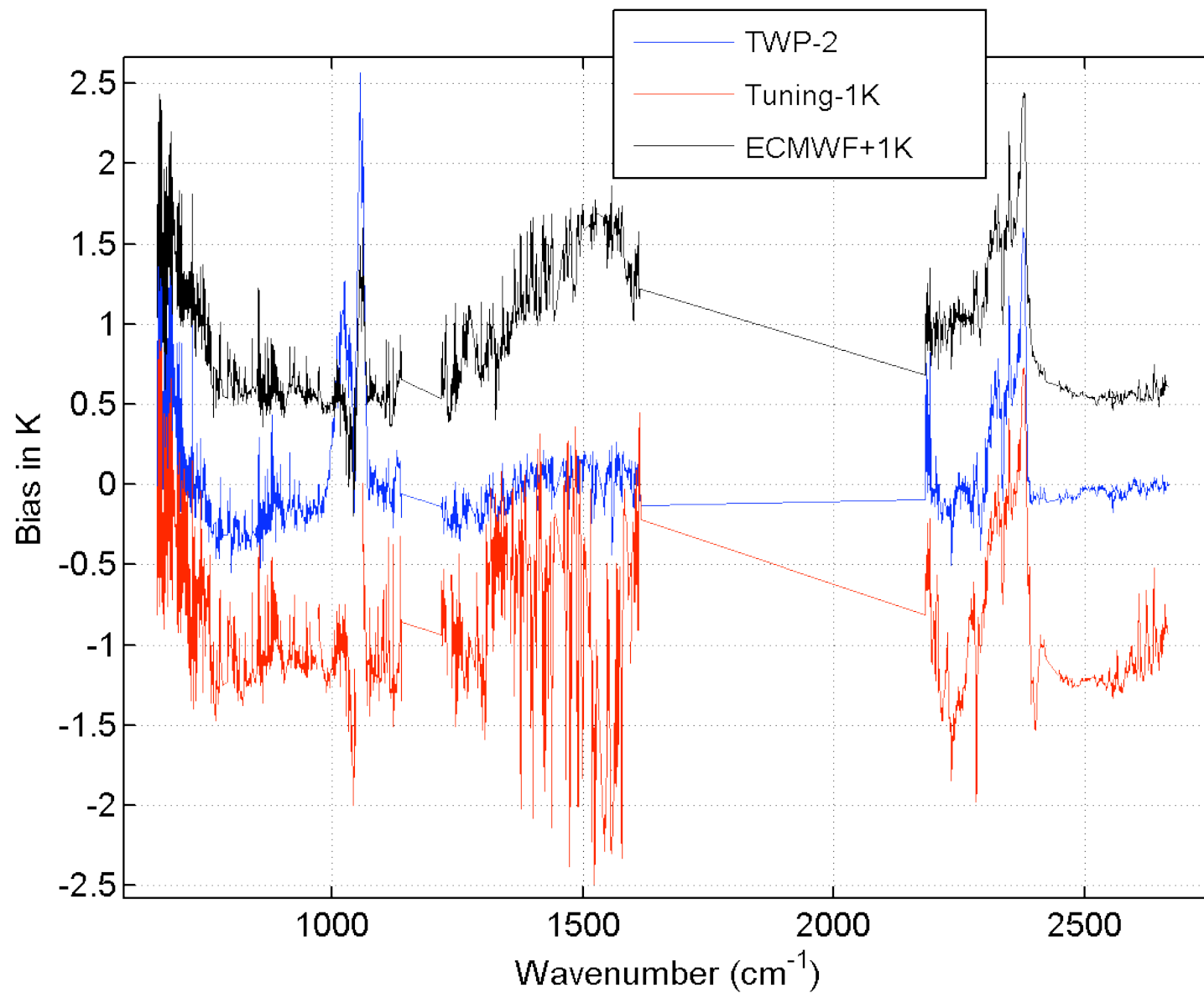
# TWP versus ECMWF



# Minnett and Voemel

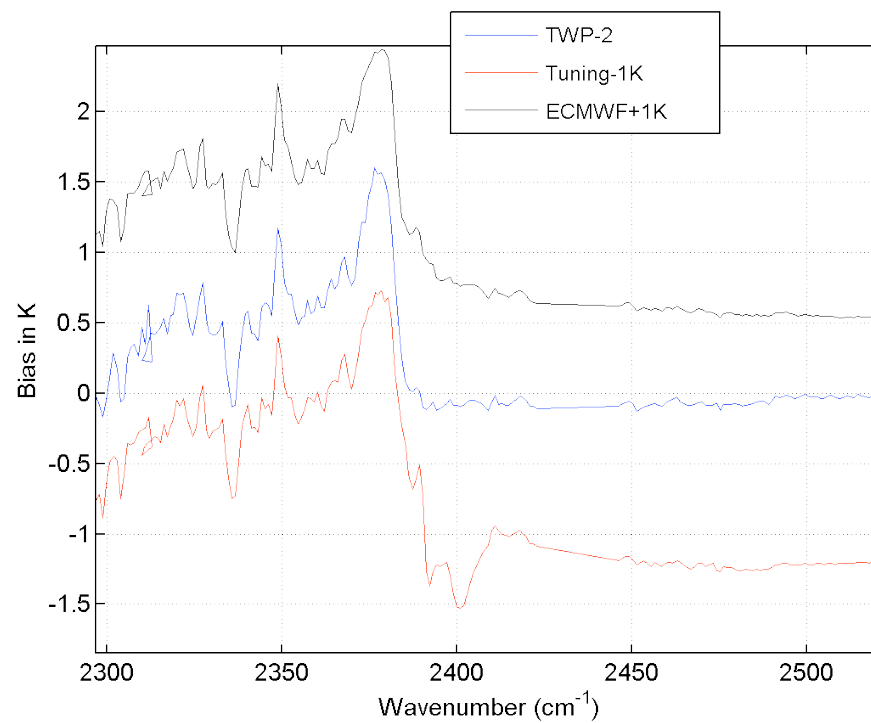
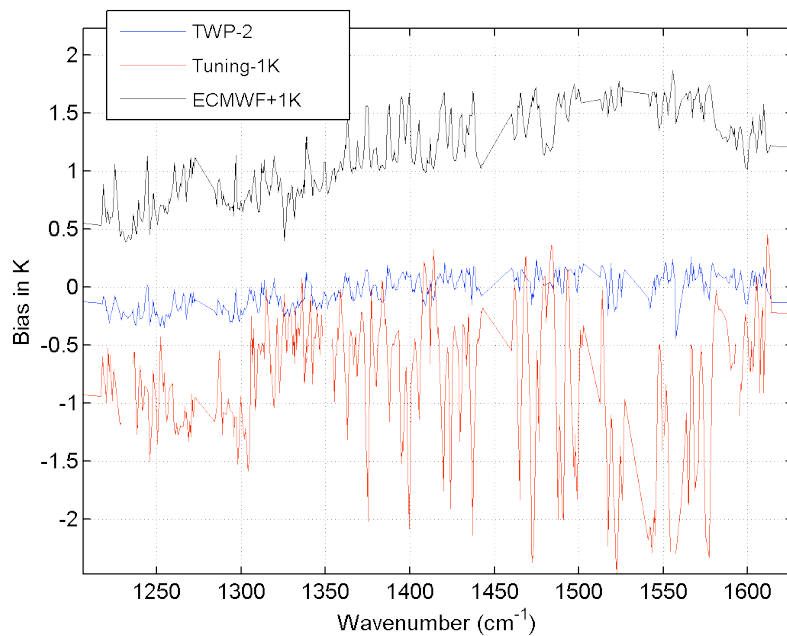
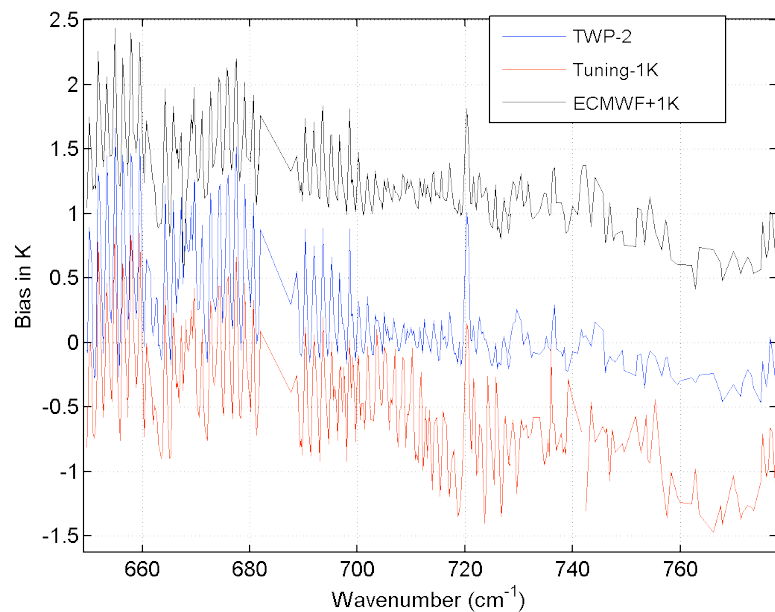


# Validation Bias vs V3.x Tuning

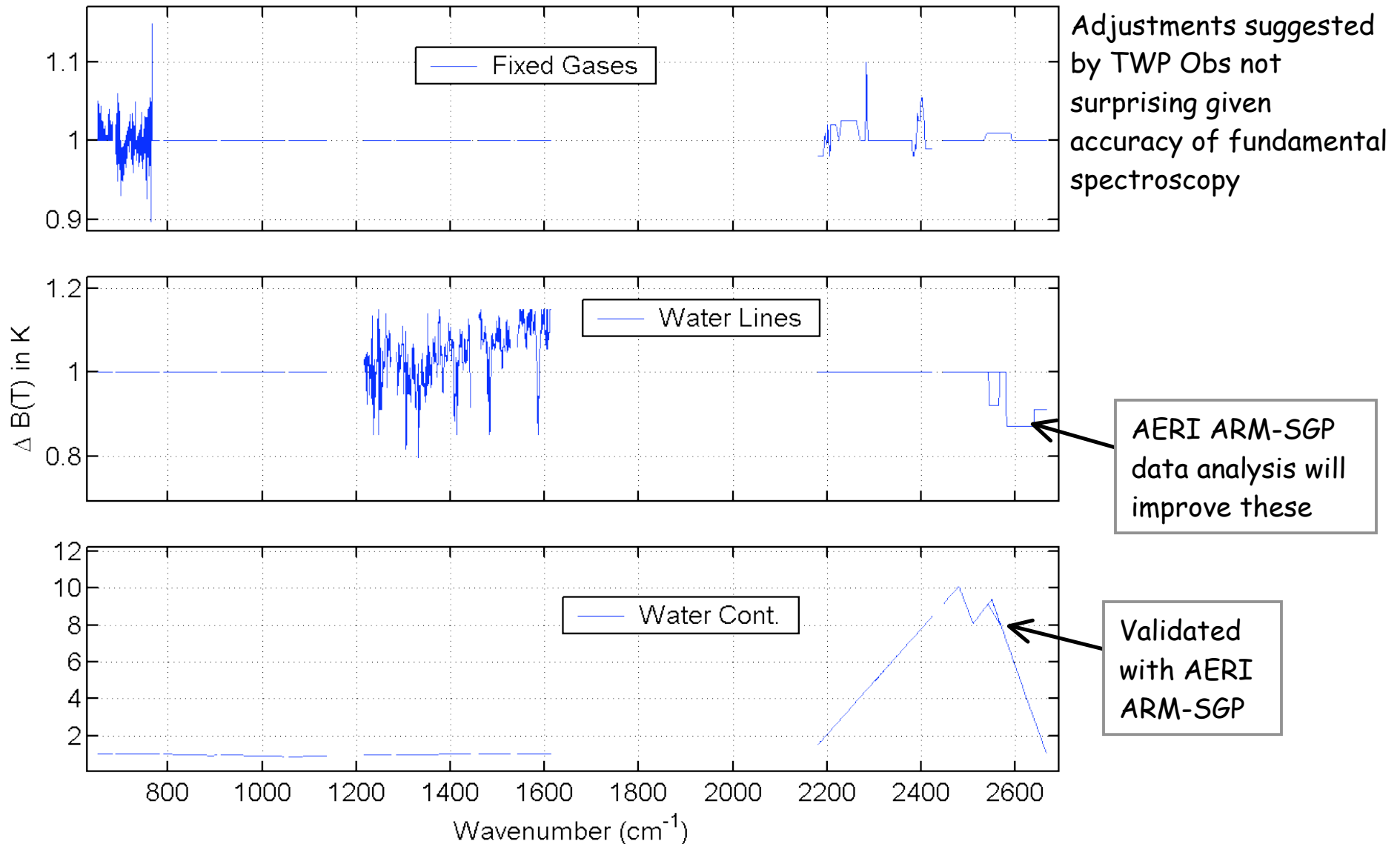




# Validation Biases vs 3.x Tuning

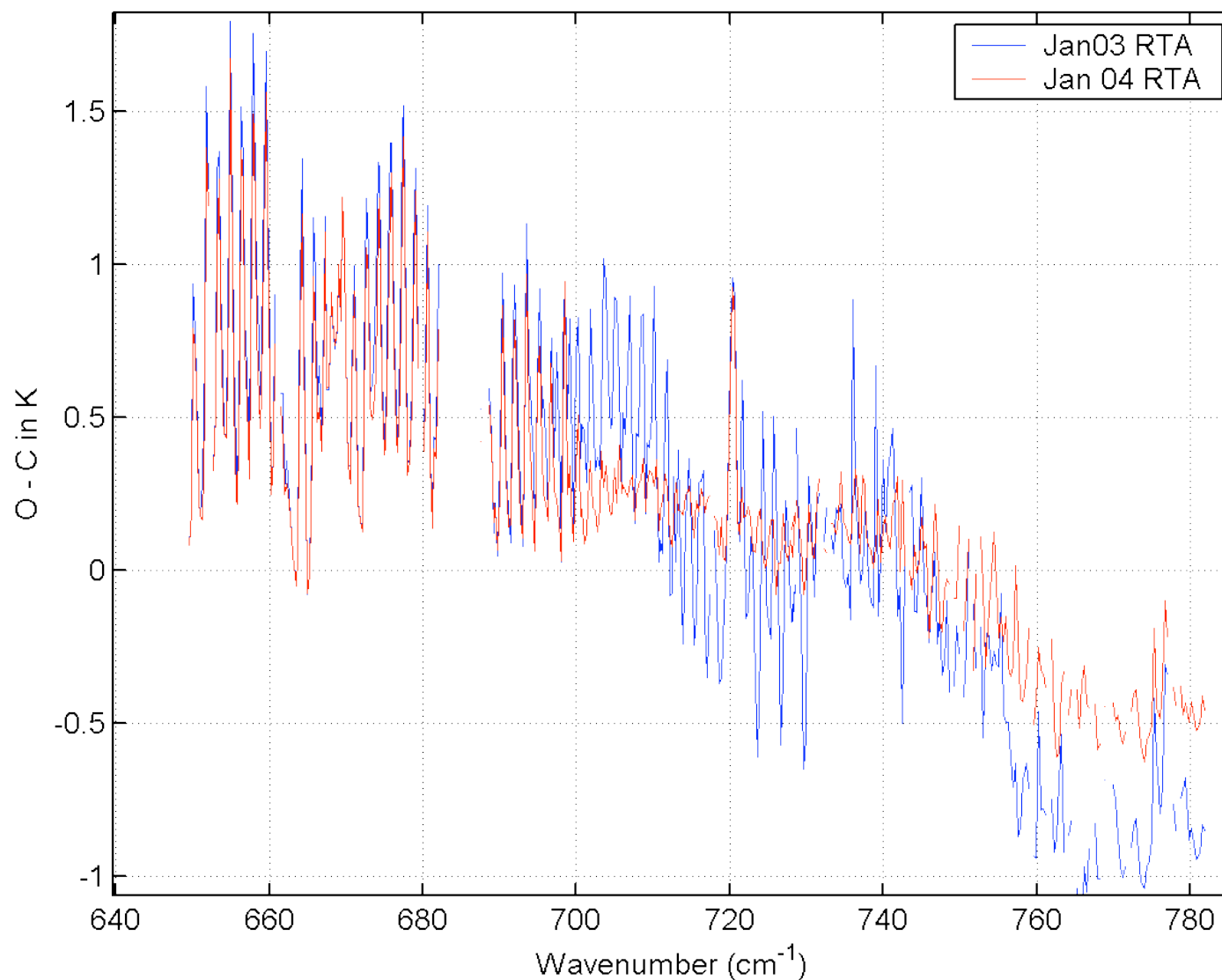


# Empirical Adjustments to RTA Transmittances



# Improvements to ECMWF Bias from ARM-TWP Adjustments

Jan03 vs Jan04 RTA: ECMWF Biases for Oct. 2002



- Freq Calibration:
  - Prototype S/W works (Matlab)
  - Fix only in L2 processing, what about L1b DAAC users?
- Fringes:
  - Can re-produce Nov 03 shifts
  - Assume we know absolute fringe positions (more modeling might help here)
- Scan Asymmetry
  - Static, but results are for clear only. Look at CC'd data.
- Non-LTE
  - Priority? We have plenty to do.
- Variable Gases
  - Use CMDL for CO<sub>2</sub> climatology? Need stratospheric climatology that doesn't exist
  - Add variable N<sub>2</sub>O to RTA. Climatology for amount?
  - CH<sub>4</sub>, handle with retrieval?
- RTA accuracy
  - With new large sonde data sets, more "tuning"? Add Miloshevich corrections and higher latitude datasets.
  - Is water band bias variability profile dependent?
  - Reflected thermal over land?